

Previous Lecture

- Cerenkov effect
- Its application to SuperKamiokande
- Start atmospheric neutrinos...

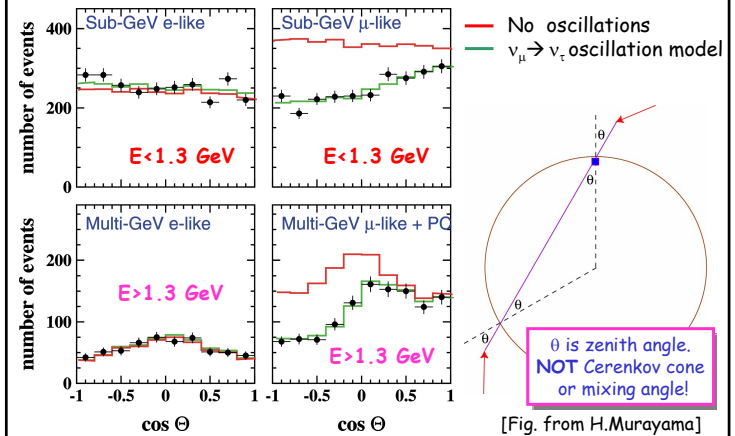
Lecture Content

- Approx. lecture content
 1. PP intro
 2. PP intro.
 3. ν props 1: strong/e.m./weak, no. neutrino generations
 4. ν props 2: lepton no., ν existence
 - Examples of decay/production
 5. Neutrino mass
 - Fermi-Kurie plot
 - Phase space kinematics/4-momentum
 6. Parity and CP violation... (why so important in lepton sector?)
 - Wu et al., ^{60}Co experiment
 7. Detection & observation
 - Liquid, solid, bubble chamber
 - "Direct" methods (DONUT)
 8. Atmospheric neutrinos
 - Cerenkov detectors
 - SuperKamiokande experiment
 9. Atmospheric neutrino data and oscillations
 - Interpretation of atmospheric ν data
 - Two-flavour neutrino oscillation formalism
 10. Solar neutrinos and SSM
 - SNO experiment and data
 - NBDB (NEMO, etc.)
 11. Implications for cosmology
 - Open vs. closed scenarios: various m_ν regions
 - ν as DM candidate?
 - Subject outlook (JPARC, MICE, Neutrino Factory, SK, SNO, KAMLAND, CHOOZ, MINOS, miniBOONE, JPARC, νF).

Today

- Atmospheric ν data and their interpretation
- Neutrino oscillation concept
- 2 flavour neutrino flavour oscillation formalism
- Generalisation to 3 flavours: MNS (Maki-Nakagawa-Sakata) mixing

Super-Kamiokande results II



Super-Kamiokande results I

- Super-K atmospheric ν data, solid evidence for oscillations (1998)
- Predicted $N(\mu)/N(e)$ ratio $\cong 2$
 - ▶ μ from $\pi^- \rightarrow \mu^- + \bar{\nu}_\mu$ and $\mu^- \rightarrow e^- + \nu_\mu + \bar{\nu}_e$
 - ▶ e only from μ^- decay
- Measured $N(\mu)/N(e)$ / predicted $N(\mu)/N(e) = 0.64 \pm 0.05$
 - ▶ Prediction includes simulation of resolution, detector thresholds, etc.
- $N(\mu)/N(e)$ increases for higher energy μ (time dilation by Lorentz γ factor)
 - ▶ More time dilation \Rightarrow fewer μ decay in atmosphere: $R = \gamma c \tau = (E/m)c\tau$
 - ▶ Remember μ from $\pi^- \rightarrow \mu^- + \bar{\nu}_\mu$ and $\mu^- \rightarrow e^- + \nu_\mu + \bar{\nu}_e$
 - ▶ e only from μ^- decay
- Clear asymmetry in upward/downward μ events
 - ▶ Shows no. of ν_μ depends on distance travelled between production (in atmosphere) and detector
 - ▶ No such effect for electrons
 - ▶ Common interpretation: $\nu_\mu \rightarrow \nu_\tau$ oscillation

"Survival" Probability

