

Previous lecture

- Quark content of mesons/baryons
 - ▶ flavour quantum number, e.g. $B^0 = b\bar{d}$
- Feynman diagrams (process description only, no calc.)
 - ▶ scattering and annihilation
- Δ^{++} decay (strong decay, with weak decay of π^+)
 - ▶ $\Delta^{++} \rightarrow p\pi^+$ (strong, $\tau \sim 10^{-23}s$)
 - ▶ $\pi^+ \rightarrow \mu^+\nu_\mu$ (weak, $\tau \sim 10^{-8}s$)

Lecture Content

■ Approx. lecture content

1. PP intro
2. PP intro.
 - Feynman diagrams; strong/e.m./weak
3. ν props 1: **baryon and lepton numbers; no. neutrino generations,**
4. ν props 2: lepton no., ν existence
 - Examples of decay/production
5. Neutrino mass
 - Fermi-Kurie plot
 - Phase space kinematics/4-momentum
6. Parity and CP violation... (why so important in lepton sector?)
 - Wu et al., ^{60}Co experiment
7. Detection & observation
 - Liquid, solid, bubble chamber
 - "Direct" methods (DONUT)
8. Solar and atmospheric neutrinos
 - Puzzle: relative abundances \neq SSM prediction
 - Two-flavour neutrino oscillation formalism
9. Neutrino oscillations and mixing
 - Possible solutions to solar/atm. ν problems
10. Current and future experiments
 - SK, SNO, KAMLAND, CHOOZ
 - MINOS, miniBOONE, ..
 - NDBD (NEMO, etc.)
 - JPARC, νF ,
11. Implications for cosmology
 - Open vs. closed scenarios. various m_ν regions
 - ν as DM candidate?
 - Subject outlook (JPARC, MICE, Neutrino Factory, ...)

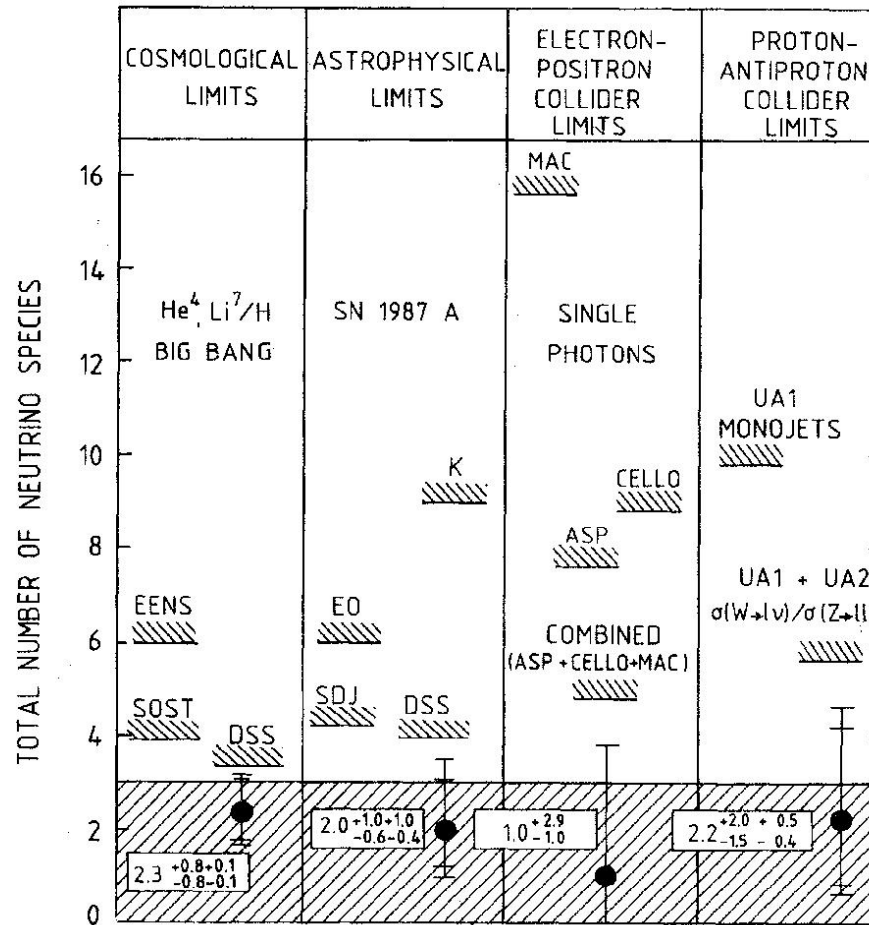
Today

- Baryon number
- Neutrino properties
 - ▶ Lepton number
 - ▶ Number of ν flavours

See also

Winter: Sect. 2.1, 2.3

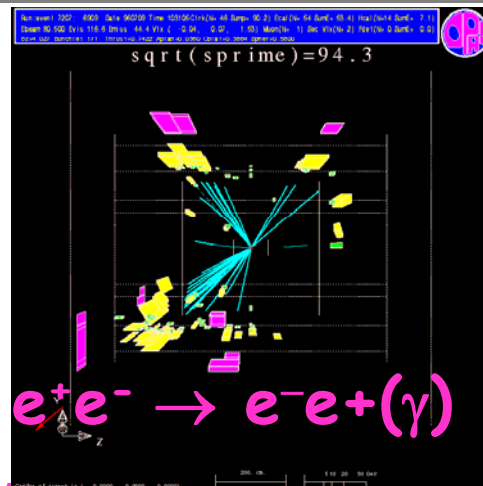
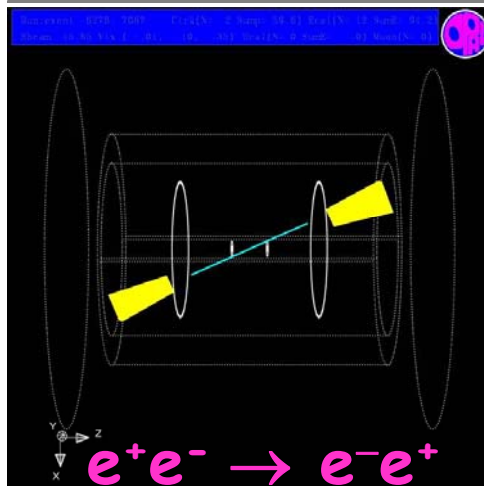
Sutton ("spaceship neutrino"), chapter 2



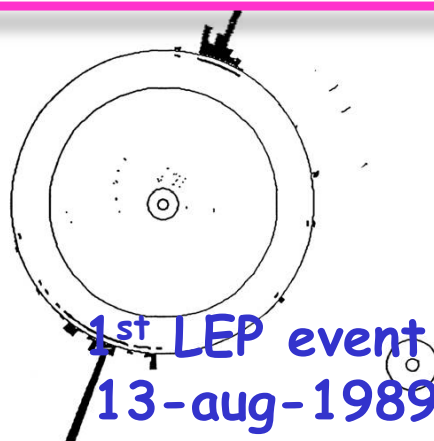
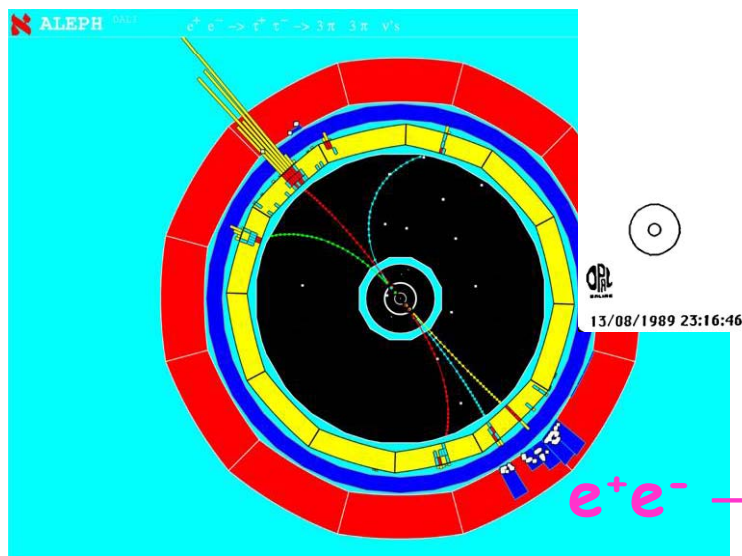
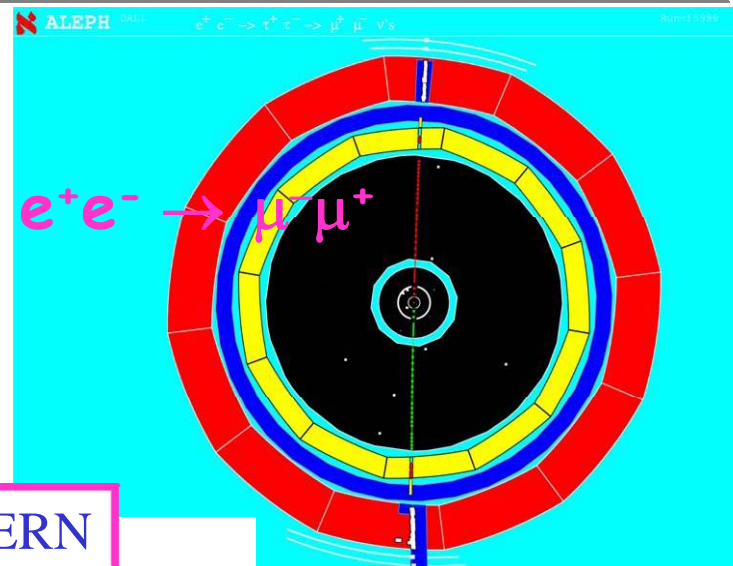
From Denegri, Sadoulet and
Spiro, Rev. Mod.
Phys. 62, 1-42 (1990)

FIG. 34. Compilation of central values and 90% C.L. upper limits on the number of neutrino flavors N_ν from cosmology, astrophysics, and particle physics. The central value of N_ν from UA1 and UA2 is for the central theoretical expectation $R_\sigma = 3.25$ and for $m_{\text{top}} = 50$ GeV. The upper limit is for the "worst case:" $R_\sigma = 3.15$ and $m_{\text{top}} = 50$ GeV. The abbreviations used for the cosmological and astrophysical limits are EENS=Ellis *et al.*, 1986; SOST=Steigman *et al.*, 1986; K=Krauss, 1987; EO=Ellis and Olive, 1987; SDJ=Shaeffer *et al.*, 1987; and DSS=this paper.

Q: How do these reactions help to measure no. ν generations?



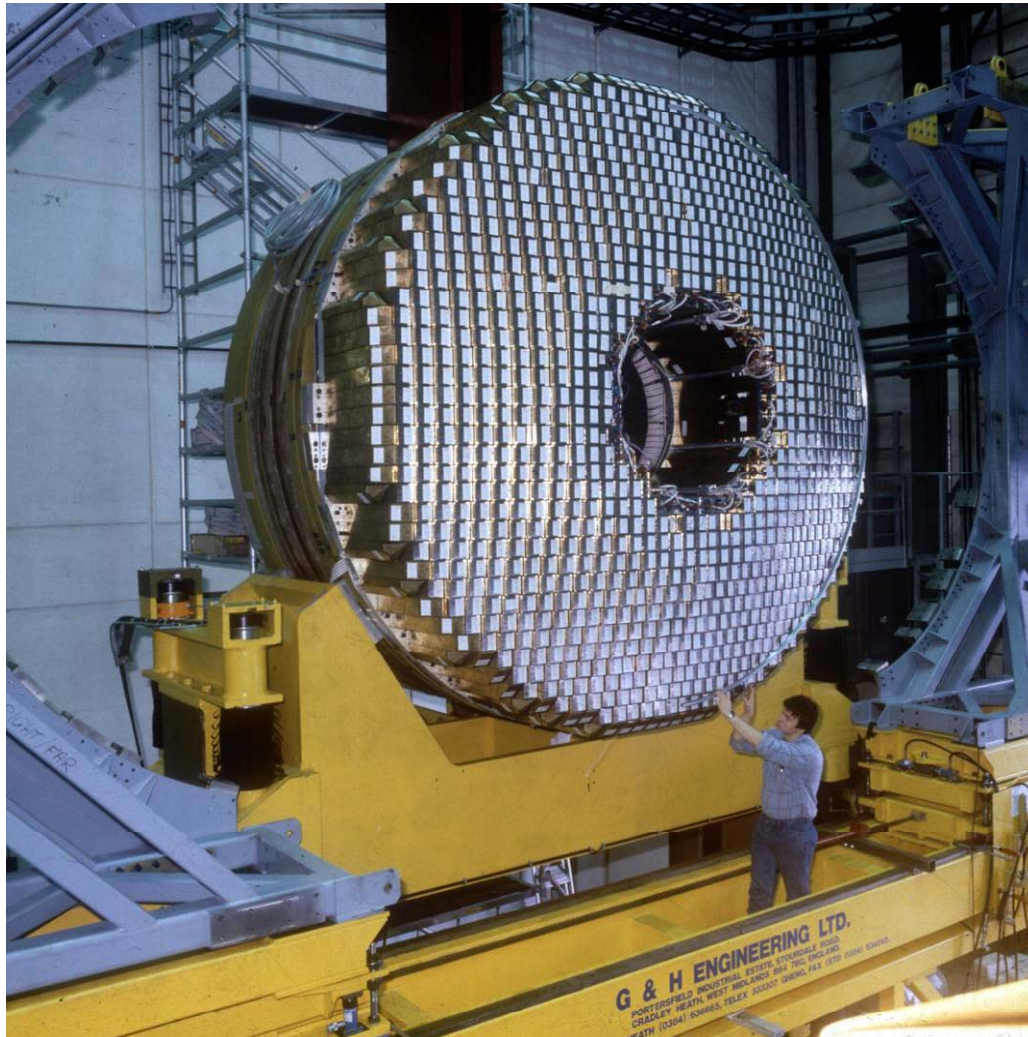
Z⁰ events from LEP1 at CERN



13-8-89
23:20



LEP expt., e.g. OPAL



Electromagnetic calorimeter
"endcap"

1132 lead glass blocks

Front face $\sim 10 \times 10 \text{ cm}^2$

Weight $\sim 25 \text{ kg/block}$

LEP: 1989-2001

■ Precise measurements $\sqrt{s} \sim m_Z$

▶ No. neutrinos

▶ Couplings, mixing angles

■ $f\bar{f}$ physics above m_Z

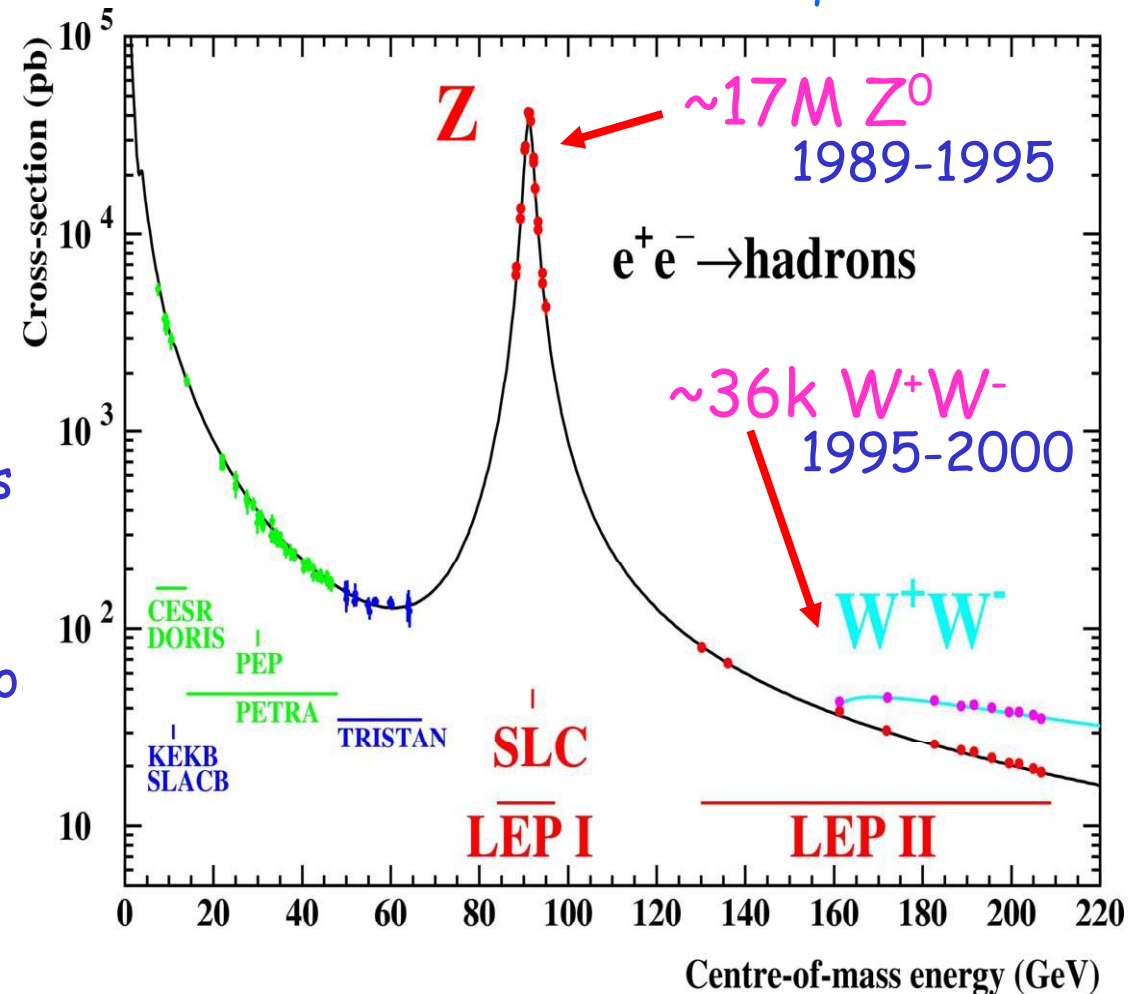
■ W^+W^- production, properties

■ Neutral boson pair production

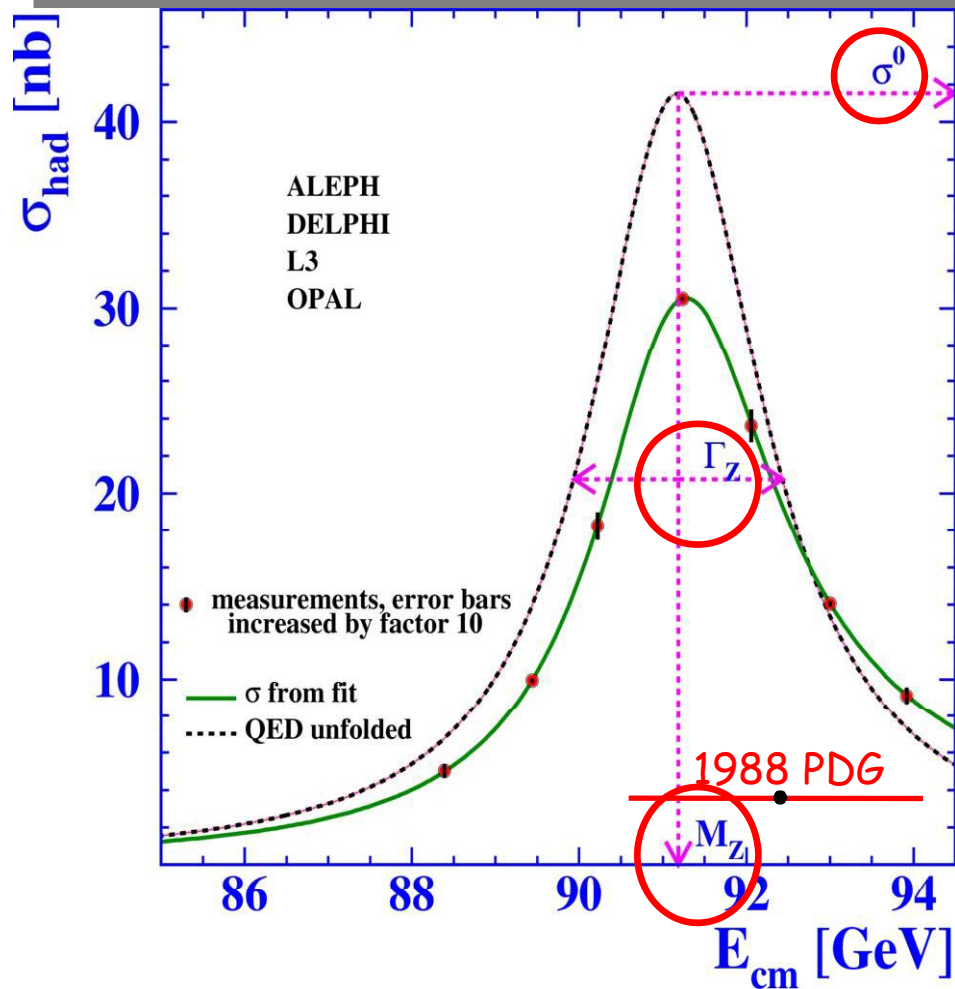
■ SM interpretations of data:

▶ Higgs mass

Data corrected for γ radiation

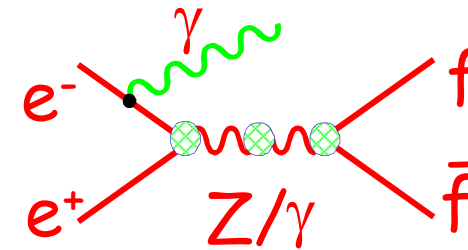


LEP Lineshape: 1989+...



3 parameters: m_Z , Γ_Z , σ_{had}^0

Observe



...EW radiative corrections

absorbed in effective couplings

Deconvolve to Born cross-section

Obtain

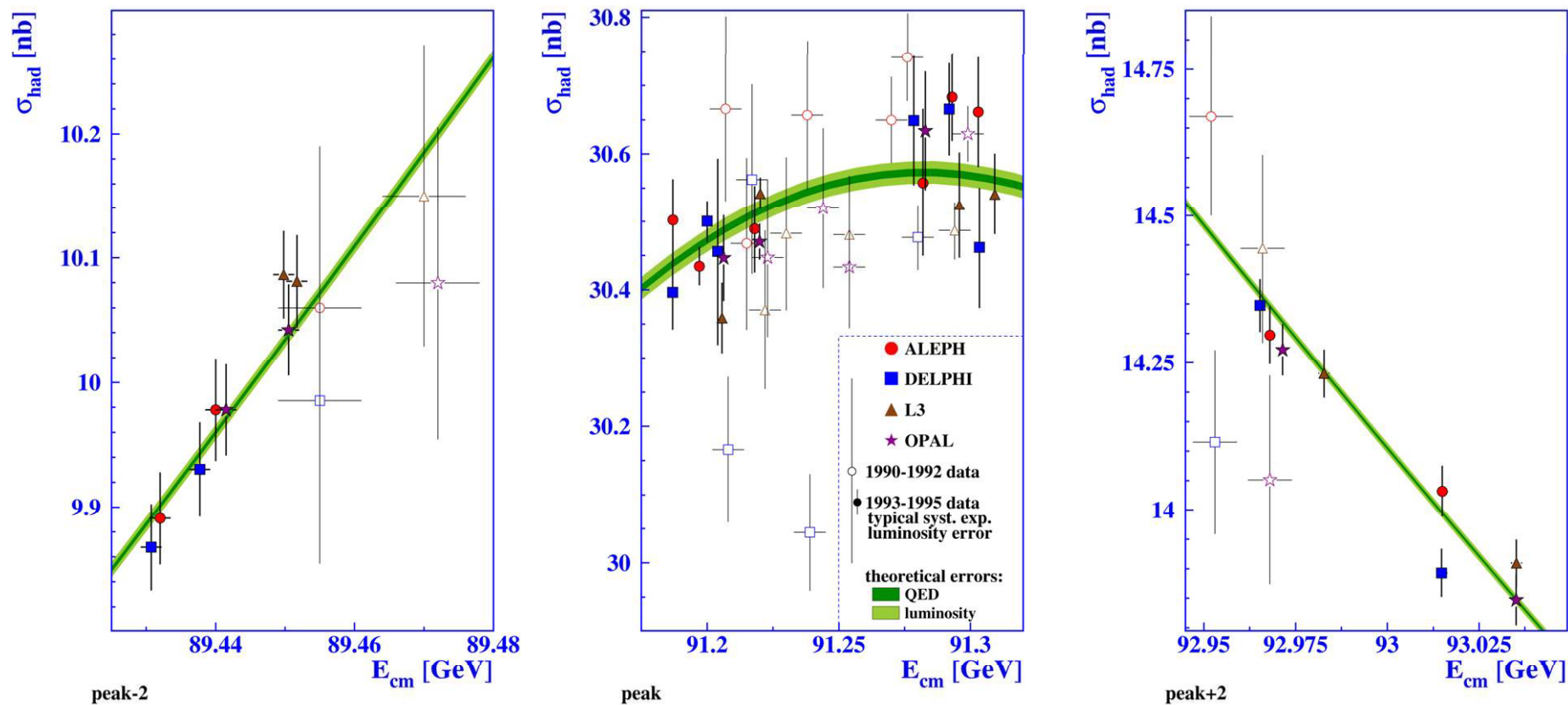
$$m_Z = 91,187.5 \pm 2.1 \text{ MeV}$$

$$\Gamma_Z = 2,495.2 \pm 2.3 \text{ MeV}$$

$$\sigma_{\text{had}}^0 = 41.540 \pm 0.037 \text{ nb}$$

Combined measurements, LepI sample

Details of LepI Cross-Section Data



All 4 LEP experiments and
years of LepI

No. of Neutrino Generations

- “Invisible width”, $\Gamma_{\text{inv}} = \Gamma_Z - \Gamma_{\text{had}} - 3 \Gamma_\ell$
- No. of generations = $\Gamma_{\text{inv}} / \Gamma_{\nu}^{\text{SM}}$ SM: $\Gamma_{\nu}^{\text{SM}} = \frac{G_F m_Z^3}{6\pi\sqrt{2}} (g_{\nu,\nu}^2 + g_{a,\nu}^2) \approx 166 \text{ MeV}$
 ▶ Measure Γ_{inv}

- Direct: measure $\sigma(e^+e^- \rightarrow \nu\bar{\nu}\gamma)$ soft γ + nothing else...challenging!

- Indirect: measure $m_Z, \Gamma_Z, R_\ell, \sigma_{\text{had}}^\circ$ $\sigma_{\text{had}}^\circ \equiv \frac{12\pi\Gamma_e\Gamma_{\text{had}}}{(m_Z\Gamma_Z)^2}$

$$\Gamma_{\text{inv}} / \Gamma_{\nu}^{\text{SM}} = \left(\frac{12\pi}{m_Z^2 \sigma_{\text{had}}^\circ} \right)^{\frac{1}{2}} - R_\ell - 3$$

$$\Rightarrow N_\nu = 2.9841 \pm 0.0083 \quad \text{for } m_\nu \leq \frac{1}{2}m_Z \sim 45 \text{ GeV}$$

- For $N_\nu = 3$, width from new Z decay modes = $-2.7 \pm 1.6 \text{ MeV}$
- Still room for heavy or sterile neutrinos