

# Previous lecture

- Neutrino Oscillations, see handout lecture 8

- ▶ Mass and weak eigenstates not the same

- Two-flavour mixing formula

$$\text{Probability}(v_e \rightarrow v_\mu) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

- ▶  $L$  (km),  $\Delta m^2$  ( $\text{eV}^2$ ),  $E$  (GeV)
- ▶ Single mixing angle,  $\theta$  (ensures matrix is unitary)
- ▶  $\theta$  a parameter of nature, fixed, no relation to angular variable
- ▶  $\theta$  defines size of mixing
- ▶  $\Delta m^2$  ( $= m_2^2 - m_1^2$ ) gives dependence of oscillations on  $L/E$
- ▶  $(E/1.27\Delta m^2)$  defines an oscillation length scale

- Generalisation to 3 neutrino flavours

- ▶ 3 mixing angles,  $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$
- ▶ 3x3 "MNS" matrix

- Direct analogy with quarks ("CKM" mixing matrix)

# Lecture Content

## ■ Approx. lecture content

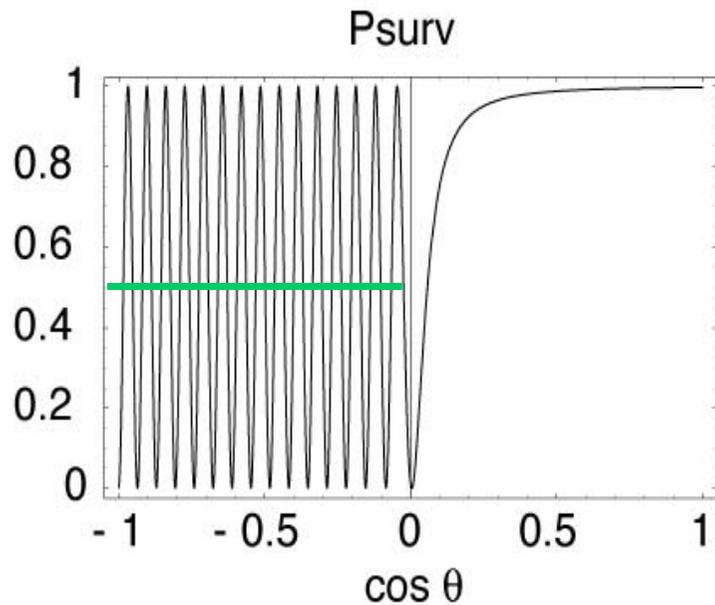
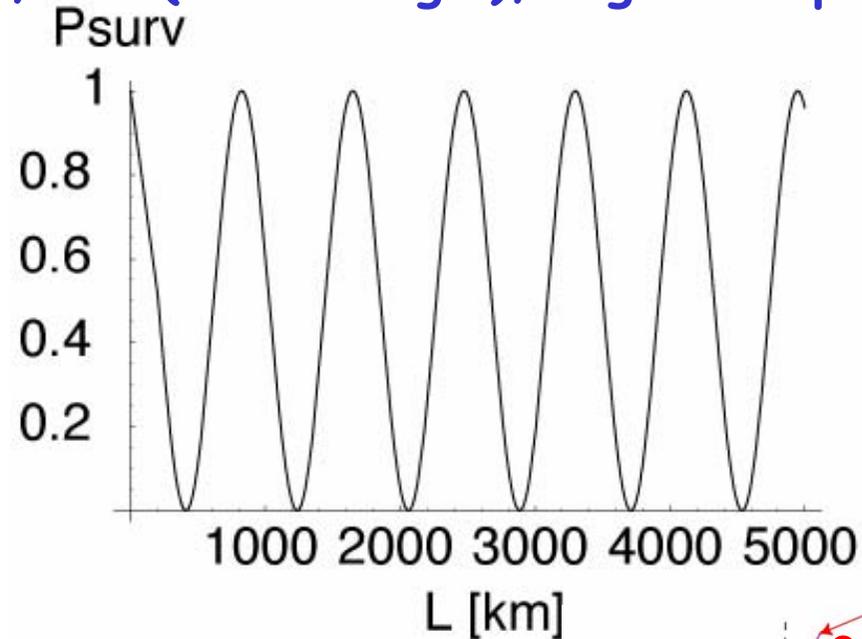
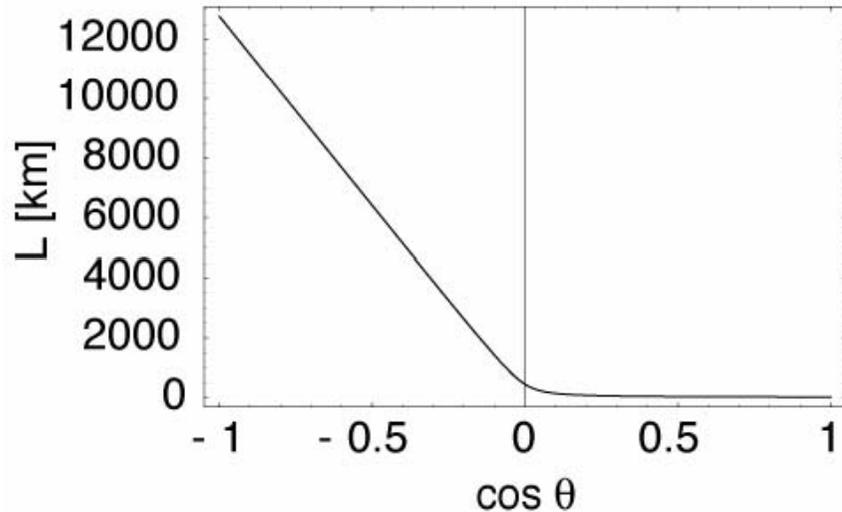
1. PP intro
2. PP intro.
3.  $\nu$  props 1: strong/e.m./weak, no. neutrino generations
4.  $\nu$  props 2: lepton no.,  $\nu$  existence  
Examples of decay/production
5. Neutrino mass  
Fermi-Kurie plot  
Phase space kinematics/4-momentum
6. Parity and CP violation... (why so important in lepton sector?)  
Wu et al.,  $^{60}\text{Co}$  experiment
7. Detection & observation  
Liquid, solid, bubble chamber  
"Direct" methods (DONUT)
8. Atmospheric neutrinos  
Cerenkov detectors  
SuperKamiokande experiment
9. Atmospheric neutrino data and oscillations  
Interpretation of atmospheric  $\nu$  data  
Two-flavour neutrino oscillation formalism
10. **Solar neutrinos and SSM**  
SNO experiment and data  
NDBD (NEMO, etc.)
11. Implications for cosmology  
Open vs. closed scenarios. various  $m_\nu$  regions  
 $\nu$  as DM candidate?  
Subject outlook (JPARC, MICE, Neutrino Factory, SK, SNO, KAMLAND, CHOOZ  
MINOS, miniBOONE, JPARC,  $\nu\text{F}$ ,

# Today

- Discussion of neutrino oscillations - See Winter 2.4.6
- "Solar neutrino problem" - See Winter, Sec. 6.1.x
- SNO
  - ▶ What it measures

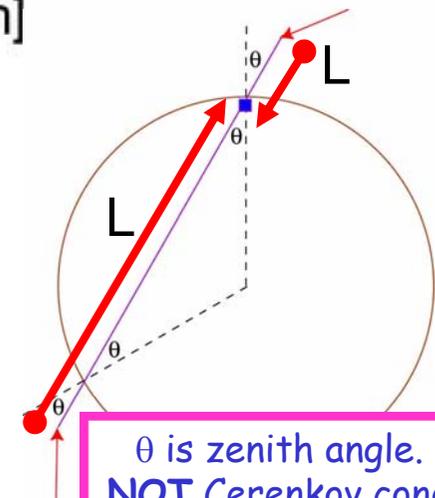
# "Survival" Probability

Distance from production,  $L$ , as  $f(\text{zenith angle})$ , e.g. in Super-K



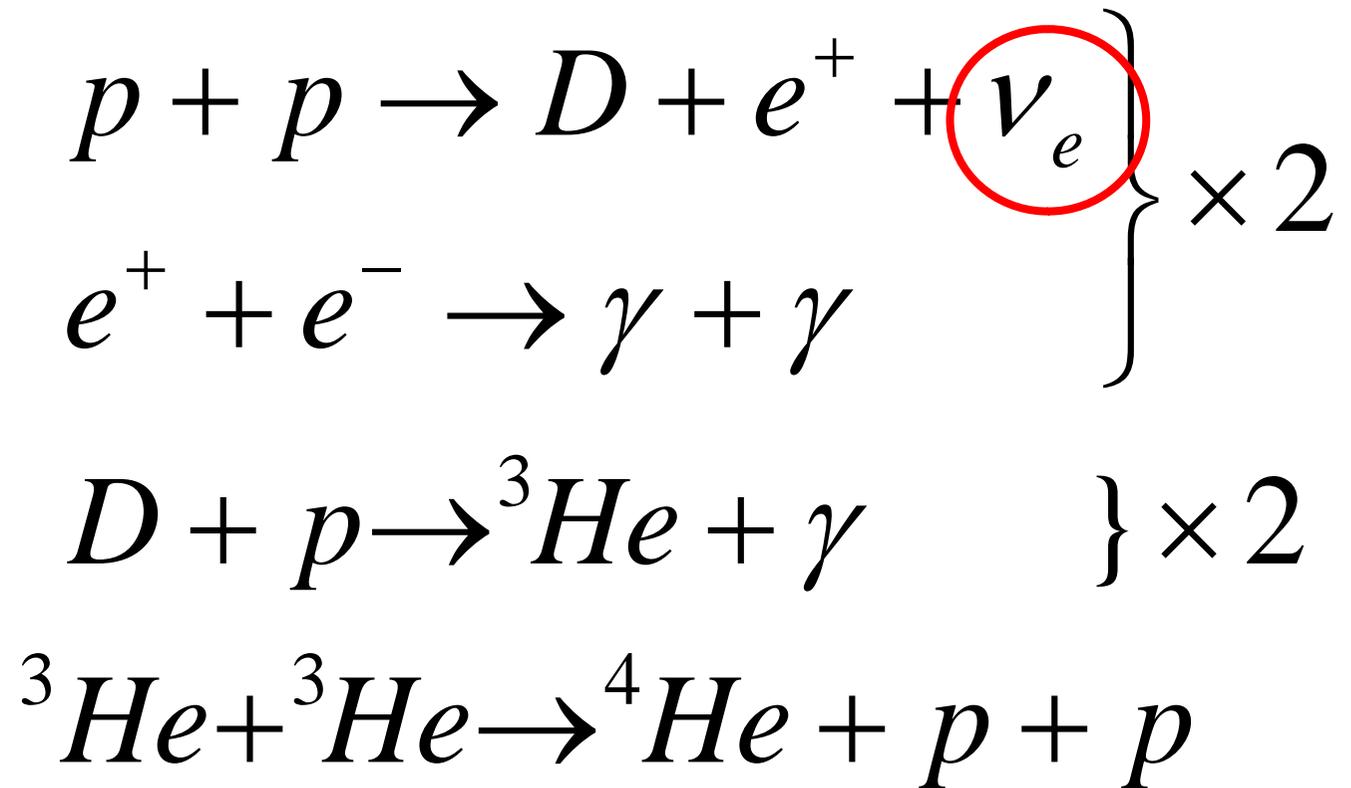
Finite angular resolution may effectively average over several oscillation cycles, net result is see  $P_{\text{surv}} \sim 0.5$

[Figs. from H.Murayama]



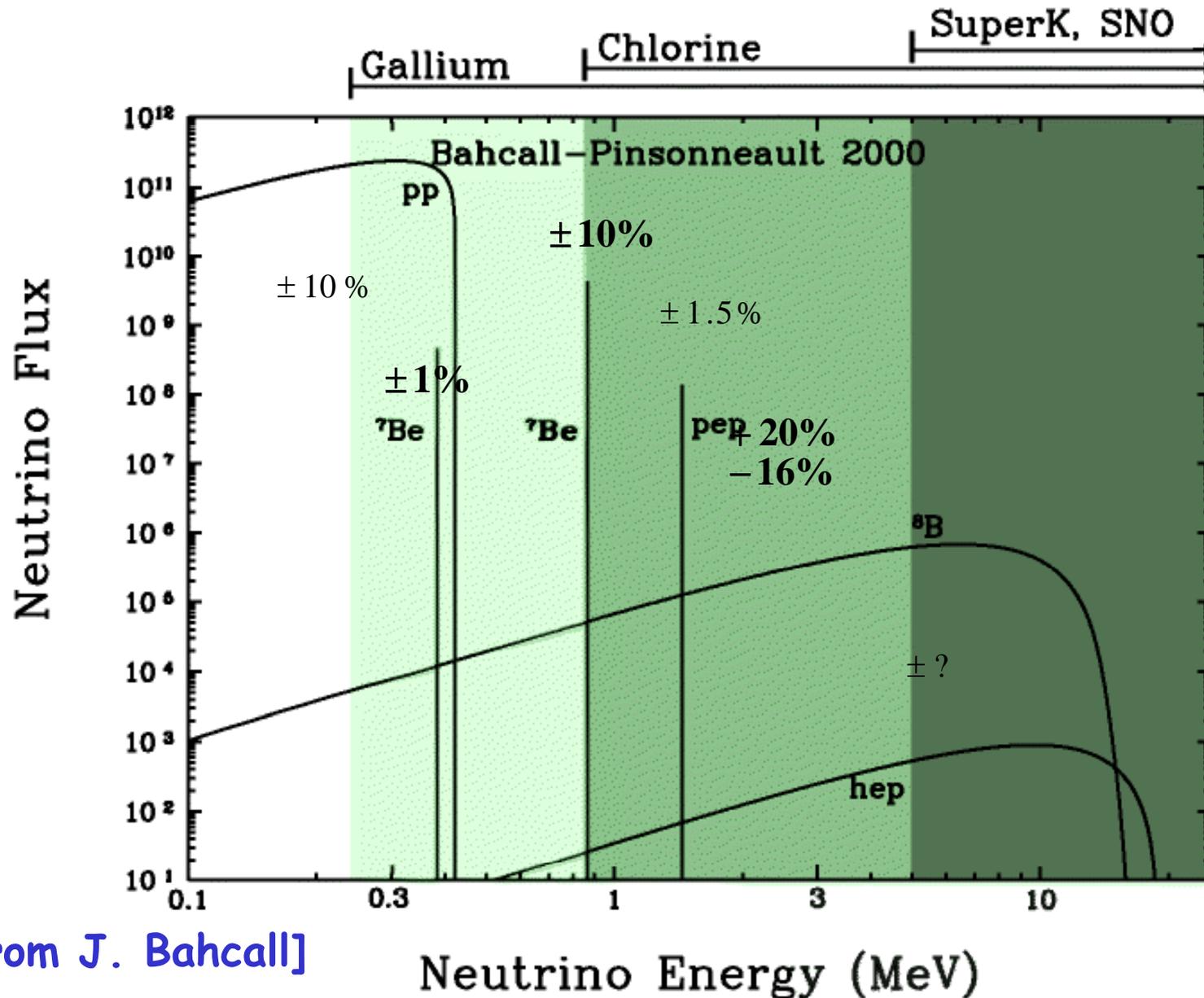
$\theta$  is zenith angle. NOT Cerenkov cone or mixing angle!

# Solar fusion: the p-p mechanism



Protons fuse, ultimately form an alpha particle, net release of 26.7 MeV/event

# $\nu_e$ Solar Neutrino Spectrum



[Fig. from J. Bahcall]

# Solar $\nu$ : Data vs. Theory

1 Solar  $\nu$  Unit =  
 $10^{-36}$  interactions/target atom/s

Total Rates: Standard Model vs. Experiment

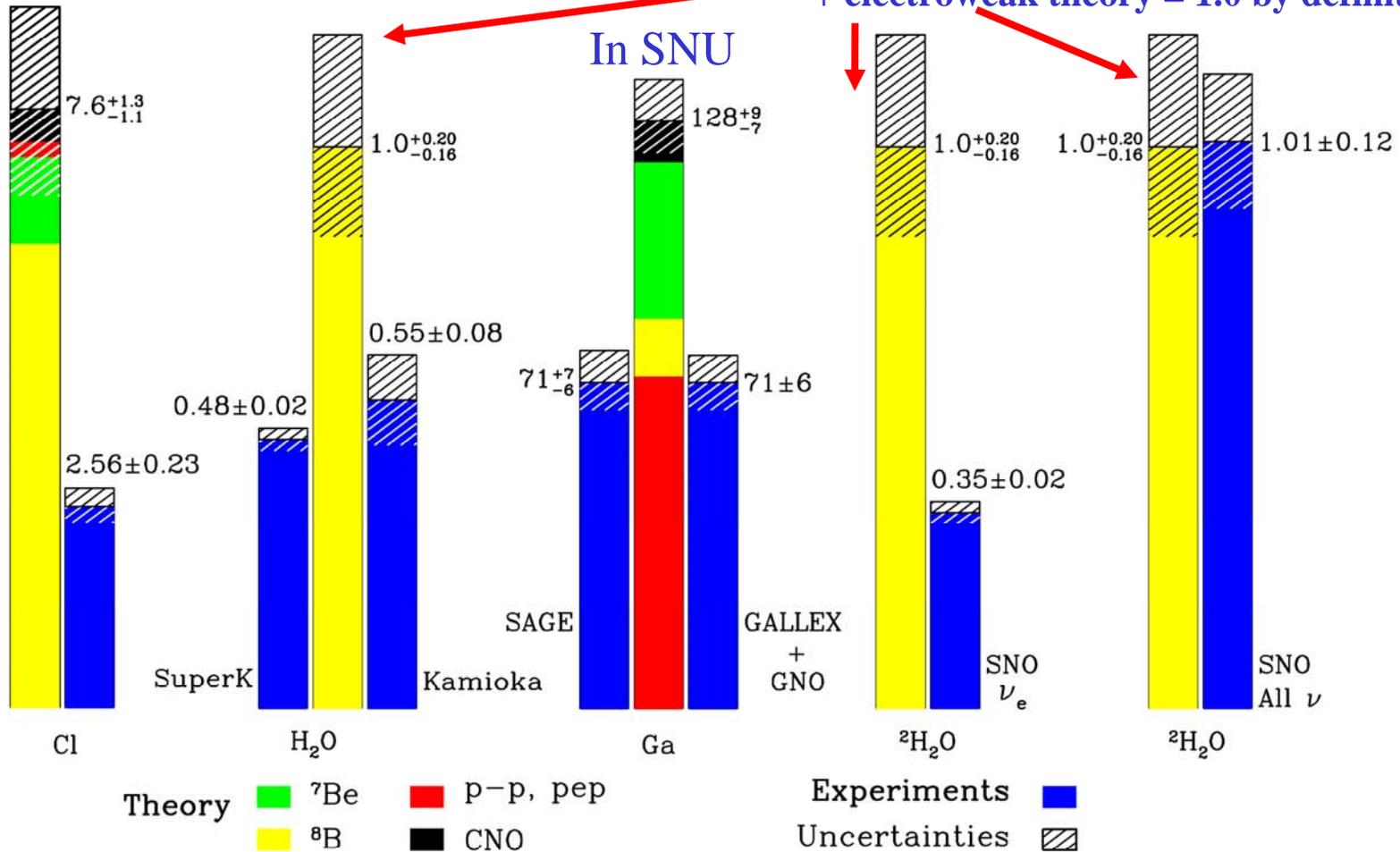
Bahcall-Pinsonneault, 2000

In units of Standard Solar Model

+ electroweak theory = 1.0 by definition

In SNU

In SNU



[Fig. from J. Bahcall]

# SNO (Sudbury Neutrino Observatory)

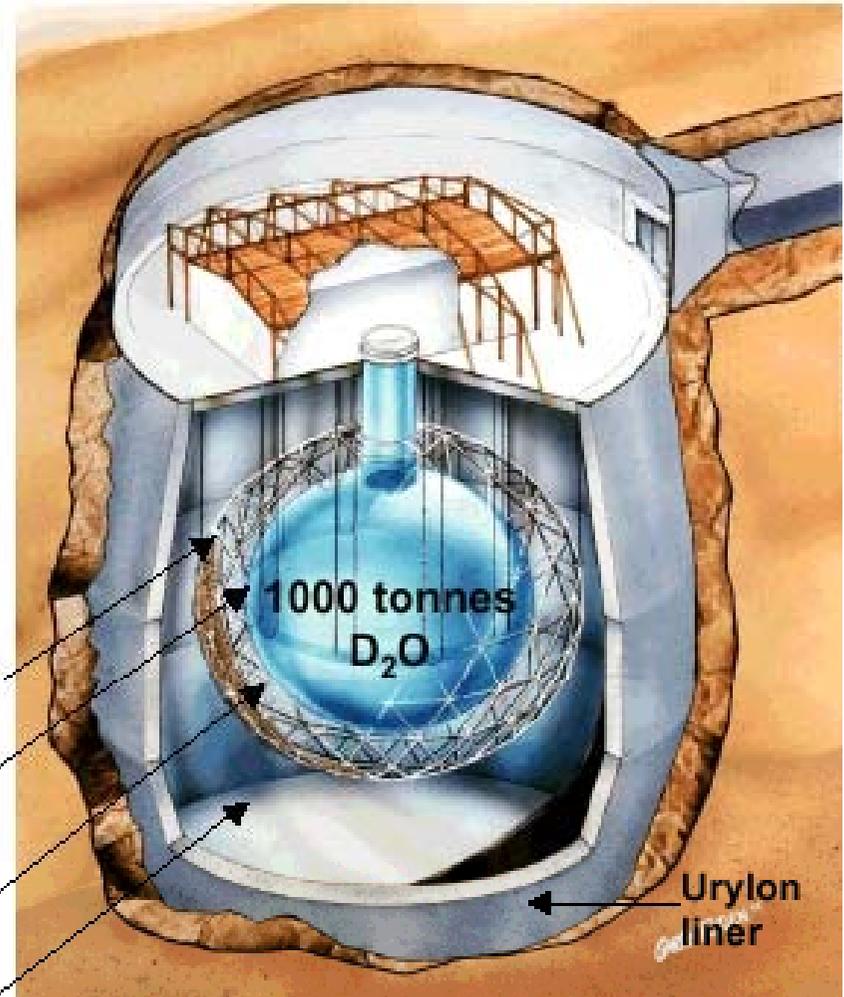


17.8m dia. PMT Support Structure  
9456 PMTs, 56% coverage

12.01m dia. acrylic vessel

1700 tonnes of inner shielding  $H_2O$

5300 tonnes of outer shielding  $H_2O$



Host: INCO Ltd., Creighton #9 mine  
Coordinates: 46°28'30"N 81°12'04"W  
Depth: 2092 m (~6010 m.w.e., ~70  $\mu$  day<sup>-1</sup>)

[ $D_2O$  on loan from Atomic Energy of Canada Ltd, value \$300M (Cnd)!]

# SNO lab

