

## Previous lecture

- Neutrino Oscillations, see handout lecture 8
  - ▶ Mass and weak eigenstates not the same
- Two-flavour mixing formula

$$\text{Probability}(\nu_e \rightarrow \nu_\mu) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

- ▶  $L$  (km),  $\Delta m^2$  ( $eV^2$ ),  $E$  (GeV)
- ▶ Single mixing angle,  $\theta$  (ensures matrix is unitary)
- ▶  $\theta$  a parameter of nature, fixed, no relation to angular variable
- ▶  $\theta$  defines size of mixing
- ▶  $\Delta m^2 (= m_2^2 - m_1^2)$  gives dependence of oscillations on  $L/E$
- ▶  $(E/1.27 \Delta m^2)$  defines an oscillation length scale
- Generalisation to 3 neutrino flavours
  - ▶ 3 mixing angles,  $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$
  - ▶ 3x3 "MNS" matrix
- Direct analogy with quarks ("CKM" mixing matrix)

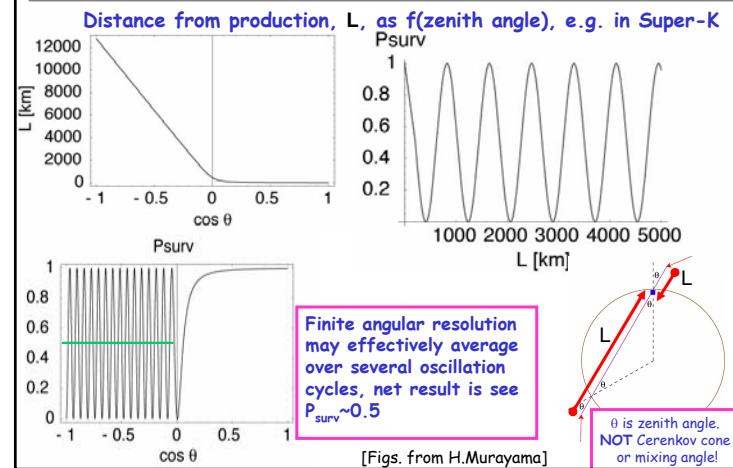
## Lecture Content

- Approx. lecture content
  1. PP intro
  2. PP intro.
  3. v props 1: strong/e.m./weak, no. neutrino generations
  4. v props 2: lepton no., v existence
    - Examples of decay/production
  5. Neutrino mass
    - Fermi-Kurie plot
    - Phase space kinematics/4-momentum
  6. Parity and CP violation... (why so important in lepton sector?)
    - Wu et al., <sup>60</sup>Co experiment
  7. Detection & observation
    - Liquid, solid, bubble chamber
    - "Direct" methods (DONUT)
  8. Atmospheric neutrinos
    - Cerenkov detectors
    - SuperKamiokande experiment
  9. Atmospheric neutrino data and oscillations
    - Interpretation of atmospheric v data
    - Two-flavour neutrino oscillation formalism
  10. Solar neutrinos and SSM
    - SNO experiment and data
    - NBDB (NEMO, etc.)
  11. Implications for cosmology
    - Open vs. closed scenarios: various  $m$ , regions
    - v as DM candidate?
    - Subject outlook (JPARC, MICE, Neutrino Factory, SK, SNO, KAMLAND, CHOOZ, MINOS, miniBOONE, JPARC, vF, ...)

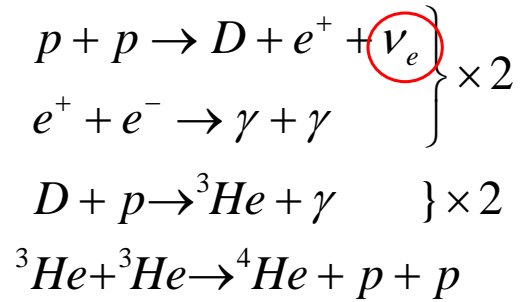
## Today

- Discussion of neutrino oscillations - See Winter 2.4.6
- "Solar neutrino problem" - See Winter, Sec. 6.1.x
- SNO
  - ▶ What it measures

## "Survival" Probability

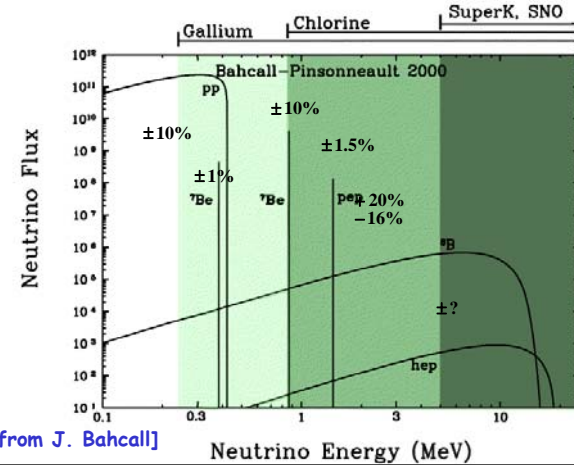


## Solar fusion: the p-p mechanism



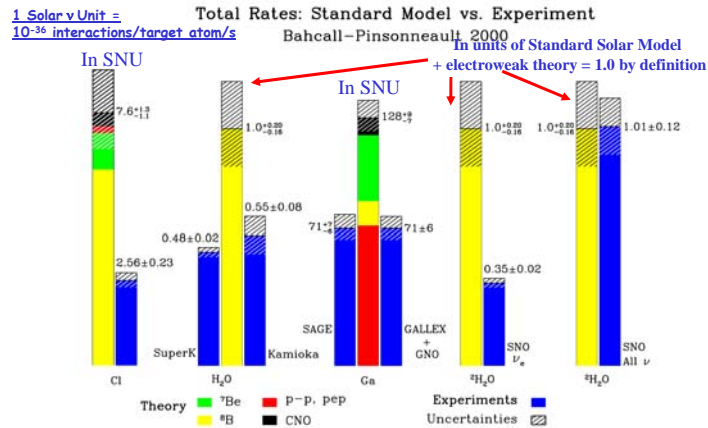
Protons fuse, ultimately form an alpha particle, net release of 26.7 MeV/event

## $\nu_e$ Solar Neutrino Spectrum

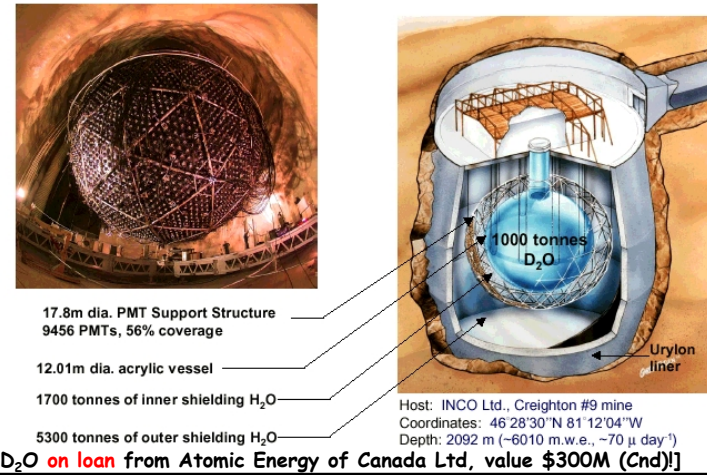


[Fig. from J. Bahcall]

## Solar $\nu$ : Data vs. Theory



## SNO (Sudbury Neutrino Observatory)



# SNO lab

