

Birmingham 11<sup>th</sup> November 2009



Conceptual Design Report: arXiv:0709.0451 (hep-ex) http://web.infn.it/superb/index.php/home

### Overview

- What is SuperB?
- Physics Case in the LHC era
- Accelerator Aspects
- Detector Design
- Current Status
- Summary

## What is SuperB?

# SuperB (In a Nutshell)

#### Site: Tor Vergata Campus (Rome II)

- Asymmetric energy e<sup>+</sup>e<sup>-</sup> collider
- Low emittance operation (like LC)
- Polarised beams
- Luminosity 10<sup>36</sup> cm<sup>-2</sup>s<sup>-1</sup>
  - 75ab<sup>-1</sup> data at the Y(4S)
  - Collect data at other vs
  - Start data taking as early as 2015
- Crab Waist technique developed to achieve these goals
- International Community



Geographical distribution of CDR signatories.



## Precision B, D and $\tau$ decay studies and spectroscopy

- New Physics in loops
  - 10 TeV reach at 75ab-1
  - Rare decays
  - ∆S CP violation measurements
- Lepton Flavour & CP Violation in τ decay
- Light Higgs searches
- Dark Matter searches

http://web.infn.it/superb/index.php/home

- Aims to constrain flavour couplings of new physics at high energy:
  - Refine understanding of nature if new physics exists at high energy.
    - We need to test the anzatz that new physics might flavour blind:
      - Case 1: trivial solution  $\rightarrow$  Reject more complicated models.
      - Case 2: non-trivial solution  $\rightarrow$  Reject flavour blind models.

Quarks and neutrinos have non-trivial couplings. e,g, the CKM matrix in the Standard Model of particle physics. How far fetched is a trivial flavour blind new physics sector?

$$J^{\mu} = \left(\overline{u} \quad \overline{c} \quad \overline{t}\right) \frac{\gamma^{\mu} \left(1 - \gamma^{5}\right)}{2} \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

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- Aims to constrain flavour couplings of new physics at high energy:
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      - Case 1: trivial solution  $\rightarrow$  Reject more complicated models.
      - Case 2: non-trivial solution  $\rightarrow$  Reject flavour blind models.
  - If the LHC doesn't find new physics: SuperB indirectly places constraints beyond the reach of the LHC and SLHC.

- The measurements to be made at SuperB fall into two categories:
  - New physics sensitive goals of the experiment
    - Some of these physics processes will be discussed in a moment: B, D, τ, Y, ....
    - This is why we want to build SuperB!
  - Standard Model calibrations (I won't talk about this)
    - This is how we validate our understanding of the detector: repeating measurements done by BaBar/ Belle and LHCb.
    - The equivalent of doing W, Z and PDF physics at ATLAS/CMS.

#### **Case studies:**

- **1. Lepton Flavour Violation**: T decay as an example of many LFV measurements possible at SuperB.
- **2. Charged Higgs:** what do we know; what will LHC tell us; what does SuperB add?
- **3. Neutral Higgs A0**: what can the flavour sector add to high  $p_T$  searches?
- 4. ΔS measurements: high mass particle interferometry.

## Physics Case in the LHC era

Why is SuperB experiment relevant when we have the energy frontier experiments and LHCb?

What is the minimum data set to make sure that we are doing something sensible?



- LHC is *not* competitive (Re: both GPDs and LHCb).
- SuperB sensitivity ~10 50× better than NP allowed branching fractions.

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- $\tau \rightarrow \mu \gamma$  upper limit can be correlated to  $\theta_{13}$  (neutrino mixing/CPV, T2K etc.) and also to  $\mu \rightarrow e \gamma$ . SUSY seasaw = CMSSM +  $3\nu_R + \tilde{\nu}$
- Complementary to flavour mixing in quarks.
- Golden modes:
  - $-\tau \rightarrow \mu \gamma$  and  $3\mu$ .
- e<sup>-</sup> beam polarization:
  - Lower background
  - Better sensitivity than competition!
- e<sup>+</sup> polarization may be used later in programme.
- CPV in  $\tau \rightarrow K_S \pi v$  at the level of ~10<sup>-5</sup>.
- Bonus:
  - Can also measure τ g-2 (polarization is crucial).
  - σ(g-2) ~2.4 ×10<sup>-6</sup> (statistically dominated error).



Use  $\mu\,\gamma/3I$  to distinguish SUSY vs. LHT.

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 $\begin{array}{ll} m_{\tilde{q}} = 300 \, GeV & \mathsf{BLUE} \\ m_{\tilde{q}} = 500 \, GeV & \mathsf{RED} \end{array}$ 



- SU(5) SUSY GUT Model (arXiv :0710.5443, Parry and Zhang).
- Model has non-trivial SUSY squark couplings.
- Current BS mixing measurement favours B(τ→μγ)>3×10<sup>-9</sup>.
- Need SuperB to probe to this sensitivity.

N.B. Different New Physics Models have different features, and different hierarchies!



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### CMSSM: LHC/SuperB complementarity



Current analysis of data prefers  $\tan\beta$ ~10.

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#### Blue = LHC:

- Will be able to measure m(A) [CP odd Higgs mass]
- Poor sensitivity to tanβ [ratio of Higgs vevs]
- •Poor sensitivity to A [coupling]

#### <u>Red=LHC+EW/Low-energy</u> <u>constraints (includes SuperB):</u>

Observable	Constraint	theo. error 0.1	
$R_{{ m BR}_{b ightarrow s\gamma}}$	$1.127\pm0.1$		
$R_{\Delta M_s}$	$0.8\pm0.2$	0.1	
$BR_{b\to\mu\mu}$	$(3.5\pm0.35) imes10^{-8}$	$2 \times 10^{-9}$	
$R_{\mathrm{BR}_{b \to \tau \nu}}$	$0.8\pm0.2$	0.1	
$\Delta a_{\mu}$	$(27.6 \pm 8.4) \times 10^{-10}$	$2.0 \times 10^{-10}$	
$M_W^{ m SUSY}$	$80.392\pm0.020\mathrm{GeV}$	0.020 GeV	
$\sin^2 \theta_W^{ m SUSY}$	$0.23153 \pm 0.00016$	0.00016	
$M_h^{\text{light}}(\text{SUSY})$	$> 114.4 \mathrm{GeV}$	$3.0\mathrm{GeV}$	

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#### Red=LHC+EW/Low-energy constraints (includes SuperB):

• Can build on the m(A) measurement to measure tanβ.

Again LHC and SuperB are complementary experiments. Each can contribute significantly to the knowledge of new physics.

### Charged Higgs: $B^{\pm} \rightarrow \tau^{\pm} \nu$

 $B^{-}$ 

h



- Within the SM, sensitive to  $f_B$  and  $|V_{ub}|$ :  $\mathcal{B}_{SM} \sim 1.6 \times 10^{-4}$ .
- $\mathcal{B}$  affected by new physics.
  - MFV models like 2HDM / MSSM.
  - Unparticles.



 $W^-H$ 

• Fully reconstruct the event (modulo v).





### **Charged Higgs**

B-factory searches competitive with LHC era: e.g. 2HDM



Charm equivalent:  $D_s^+ \rightarrow \mu^+ \nu, \tau^+ \nu$ 

## **Charged Higgs**



- Multi TeV search capability for large tanβ.
- Includes SM uncertainty ~20% from V<sub>ub</sub> and f<sub>B</sub>.
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### $\Delta S$ measurements

- β=(21.1±0.9)° from Charmonium decays.
- Look in many different b→s and b→d decays for sin2β deviations from the SM:

 $\Delta S_{\rm NP} = S_{eff} - S_{c\overline{c}s} - \Delta S_{\rm SM}$ 

• The golden channel is:



 Deviations would be from high mass particles in loops: Η, χ, ... May 2009
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### $\Delta S$ measurements

- The SM uncertainty is strongly mode dependent.
- Golden modes have to be well measured and theoretically clean.
- Prefer to also have robust constraints from more than one theoretical approach.
- Precision measurements of the reference Charmonium decay also have a small SM uncertainty.



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### $\Delta S$ measurements

- We were reminded that we should be careful with what we compare:
  - NP could affect  $c\bar{c}s sin 2\beta$ .
- 1) Predict sin2 $\beta$  from indirect constraints.  $[\sin(2\beta)]_{\text{no }V_{ub}}^{\text{prediction}} = 0.87 \pm 0.09$ .
- 2) Compare to ccs measurement.  $[\sin 2\beta]_{c\bar{cs}} = 0.672 \pm 0.023$
- 3) Compare to clean penguin measurements.  $[\sin 2\beta]_{b \rightarrow s-penguin}^{clean} = 0.58 \pm 0.06$

(or the average of the two) Are these 2.1-2.7σ hints for new physics?

Lunghi and Soni, Phys.Lett.B**666** 162-165 (2008). Buras and Guadagnoli Phys Rev D **78** 033005 (2008).

 Can theory error be reduced for other modes?



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#### Standard Model measurements.

Observable	$B$ Factories (2 $\rm ab^{-1})$	and the second se
$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)
$\cos(2\beta) (J/\psi K^{*0})$	0.30	0.05
$sin(2\beta)$ (Dh <sup>0</sup> )	0.10	0.02
$\cos(2\beta)$ (Dh <sup>0</sup> )	0.20	0.04
$S(J/\psi \pi^{0})$	0.10	0.02
$S(D^{+}D^{-})$	0.20	0.03
$S(\phi K^0)$	0.13	0.02 (*)
$S(\eta' K^0)$	0.05	0.01 (*)
$S(K_{s}^{0}K_{s}^{0}K_{s}^{0})$	0.15	0.02 (*)
$S(K_{8}^{0}\pi^{0})$	0.15	0.02 (*)
$S(\omega K_s^0)$	0.17	0.03 (+)
$S(f_0K_s^0)$	0.12	0.02 (*)
$\gamma (B \rightarrow DK, D \rightarrow CP \text{ eigensta})$	tes) $\sim 15^{\circ}$	$2.5^{\circ}$
$\gamma (B \rightarrow DK, D \rightarrow \text{suppressed})$	states) $\sim 12^{\circ}$	$2.0^{\circ}$
$\gamma (B \rightarrow DK, D \rightarrow \text{multibody s})$	tates) $\sim 9^{\circ}$	$1.5^{\circ}$
$\gamma (B \rightarrow DK, \text{ combined})$	$\sim 6^{\circ}$	$1-2^{\circ}$
$\alpha (B \rightarrow \pi \pi)$	$\sim 16^{\circ}$	3°
$\alpha (B \rightarrow \rho \rho)$	$\sim 7^{\circ}$	1-2° (*)
$\alpha (B \rightarrow \rho \pi)$ $\alpha (B \rightarrow \rho \pi)$	$\sim 12^{\circ}$	2°
$\alpha$ (combined)	$\sim 6^{\circ}$	$1-2^{\circ}$ (*)
$2\beta + \gamma (D^{(*)\pm}\pi^{\pm}, D^{\pm}K_{s}^{0}\pi^{\pm})$	20°	5°
V <sub>ab</sub>   (exclusive)	4% (*)	1.0% (*)
$ V_{cb} $ (inclusive)	1% (*)	0.5% (*)
V <sub>ub</sub>   (exclusive)	8% (+)	3.0% (*)
Vub  (inclusive)	8% (*)	2.0% (*)
		191 1910 - 1910 - 1910 - 1910
$B(B \rightarrow \tau \nu)$	20%	4% (†)
$B(B \rightarrow \mu\nu)$	visible	5%
$\mathcal{B}(B \rightarrow D\tau\nu)$	10%	2%
$B(B \rightarrow \rho \gamma)$	15%	3% (†)
$\mathcal{B}(B \rightarrow \omega \gamma)$	30%	5%
$A_{CP}(B \rightarrow K^*\gamma)$	0.007 (†)	0.004 († *)
$A_{CP}(B \rightarrow \rho \gamma)$	$\sim 0.20$	0.05
$A_{CP}(b \rightarrow s\gamma)$	0.012 (†)	0.004 (†)
$A_{CP}(b \rightarrow (s + d)\gamma)$	0.03	0.006 (†)
$S(K_{\pi}^{0}\pi^{0}\gamma)$	0.15	0.02 (*)
$S(\rho^0\gamma)$	possible	0.10
1999	0.27	
$A_{CP}(B \rightarrow K^*\ell\ell)$	7%	1%
$A^{FB}(B \rightarrow K^*\ell\ell)s_0$	25%	9%
$A^{FB}(B \rightarrow X_s \ell \ell) s_0$	35%	5%
$\mathcal{B}(B \rightarrow K \nu \overline{\nu})$	visible	20%
$\mathcal{B}(B \rightarrow \pi \nu \bar{\nu})$		possible
man - mont		honoting

**B Physics at Y(4S)** Rare Charm Decays: 1 month at  $\psi(3770)$ Channel Sensitivity

onenner	Senerei (10)
$D^0 \rightarrow e^+ e^-, \ D^0 \rightarrow \mu^+ \mu^-$	$1 \times 10^{-8}$
$D^0 \rightarrow \pi^0 e^+ e^-, \ D^0 \rightarrow \pi^0 \mu^+ \mu^-$	$2 \times 10^{-8}$
$D^0 \to \eta e^+ e^-, \ D^0 \to \eta \mu^+ \mu^-$	$3  imes 10^{-8}$
$D^0 \rightarrow K^0_s e^+ e^-, D^0 \rightarrow K^0_s \mu^+ \mu^-$	$3 \times 10^{-8}$
$D^+ \rightarrow \pi^+ e^+ e^-, \ D^+ \rightarrow \pi^+ \mu^+ \mu^-$	$1 \times 10^{-8}$
$D^0 \rightarrow e^{\pm} \mu^{\mp}$	$1 \times 10^{-8}$
$D^+ \rightarrow \pi^+ e^{\pm} \mu^{\mp}$	$1 \times 10^{-8}$
$D^0 \rightarrow \pi^0 e^{\pm} \mu^{\mp}$	$2 \times 10^{-8}$
$D^0 \rightarrow \eta e^{\pm} \mu^{\mp}$	$3 \times 10^{-8}$
$D^0 \rightarrow K_s^0 e^{\pm} \mu^{\mp}$	$3 \times 10^{-8}$
$D^+ \rightarrow \pi^- e^+ e^+  D^+ \rightarrow K^- e^+ e^+$	$1 \times 10^{-8}$

$D^+ \rightarrow \pi^- e^+ e^+, \ D^+ \rightarrow K^- e^+ e^+$	$1 \times 10^{-8}$
$D^+ \rightarrow \pi^- \mu^+ \mu^+, \ D^+ \rightarrow K^- \mu^+ \mu^+$	$1 \times 10^{-8}$
$D^+ \to \pi^- e^\pm \mu^\mp,  D^+ \to K^- e^\pm \mu^\mp$	$1 \times 10^{-8}$

Process	Sensitivity
$B(\tau \rightarrow \mu \gamma)$	$2 \times 10^{-9}$
$\mathcal{B}(\tau \rightarrow e \gamma)$	$2  imes 10^{-9}$
$B(\tau \rightarrow \mu \mu \mu)$	$2 \times 10^{-10}$
$\mathcal{B}(\tau \rightarrow eee)$	$2\times 10^{-10}$
$B(\tau \rightarrow \mu \eta)$	$4\times 10^{-10}$
$B(\tau \rightarrow e\eta)$	$6\times 10^{-10}$
$B(\tau \rightarrow \ell K_s^0)$	$2 \times 10^{-10}$

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Mode	Observable	$\Upsilon(4S)$	$\psi(3770)$	$\frown$
Mode	Observable	$(75 \text{ ab}^{-1})$	$(300 \text{ fb}^{-1})$	$\mathcal{L}$
$D^0 \rightarrow K^+ \pi^-$	$x^{\prime 2}$	(15  ab) $3 \times 10^{-5}$	(300 10 )	ha
$D \rightarrow K \cdot \pi$				
-0	y'	$7 \times 10^{-4}$		rm
$D^0 \rightarrow K^+ K^-$	$y_{CP}$	$5 \times 10^{-4}$		
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	x	$4.9\times10^{-4}$		$\leq$
	y	$3.5\times10^{-4}$		Mixi
	q/p	$3  imes 10^{-2}$		in
	$\phi$	$2^{\circ}$		ы С
$\psi(3770) \rightarrow D^0 \overline{D}^0$	$x^2$		$(1-2) \times 10^{-5}$	
	y		$(1-2) \times 10^{-3}$	
	$\cos \delta$		(0.01 - 0.02)	

See CDR and Valencia report for details of the SM measurements and other possible NP searches.

B Physics at Y(5S)					
Observable	Error with 1 $ab^{-1}$	Error with $30 \text{ ab}^{-1}$			
$\Delta\Gamma$	$0.16 \text{ ps}^{-1}$	$0.03 \ {\rm ps}^{-1}$			
Г	$0.07 \ {\rm ps}^{-1}$	$0.01 \ {\rm ps^{-1}}$			
$\beta_s$ from angular analysis	$20^{\circ}$	$8^{\circ}$			
$A_{SL}^s$	0.006	0.004			
$A_{\rm CH}$	0.004	0.004			
${\cal B}(B_s\to \mu^+\mu^-)$	-	$<8\times10^{-9}$			
$ V_{td}/V_{ts} $	0.08	0.017			
$\mathcal{B}(B_s \to \gamma \gamma)$	38%	7%			
$\beta_s$ from $J/\psi\phi$	$10^{\circ}$	$3^{\circ}$			
$\beta_s$ from $B_s \to K^0 \bar{K}^0$	$24^{\circ}$	11°			

### Golden Matrix

#### No one smoking gun... rather a 'golden matrix'.

	$H^+$ high tan $\beta$	1.00	Non-MFV	NP Z-penguins	Right-handed currents	LTH	SUSY
$\mathcal{B}(B \rightarrow X_s \gamma)$		L	M		M		
$\mathcal{A}_{CP}(B \to X_s \gamma)$			L		M		
$\mathcal{B}(B \to \tau \nu)$	L-CKM						
$\mathcal{B}(B \to X_s \ell \ell)$			М	M	M		
$\mathcal{B}(B \to K \nu \overline{\nu})$			M	L			
$S_{K_S\pi^0\gamma}$					L		
The angle $\beta$ ( $\Delta S$ )			L-CKM		L		
$\tau \rightarrow \mu \gamma$							L
$\tau \rightarrow \mu \mu \mu$						L	

... + other generic models L = Large effect. M = Measureable effect. CKM= Precision CKM (from SuperB) required.

... + charm + spectroscopy (DM /Light Higgs etc).

 Need to measure all observables in order to select/eliminate new physics scenarios!

Mode	Sensitivity			
	Current	$10 \text{ ab}^{-1}$	$75 \text{ ab}^{-1}$	
$\mathcal{B}(B \to X_s \gamma)$	7%	5%	3%	
$A_{CP}(B  o X_s \gamma)$	0.037	0.01	0.004 - 0.005	
$\mathcal{B}(B^+ \to \tau^+ \nu)$	30%	10%	3 - 4%	
$\mathcal{B}(B^+ \to \mu^+ \nu)$	X	20%	5 - 6%	
$\mathcal{B}(B \to X_s l^+ l^-)$	23%	15%	4-6%	
$A_{\rm FB}(B \to X_s l^+ l^-)_{s_0}$	X	30%	4-6%	
$\mathcal{B}(B \to K \nu \overline{\nu})$	X	X	16 - 20%	
$S(K_S^0\pi^0\gamma)$	0.24	0.08	0.02 - 0.03	

#### The golden modes

• will be measured by SuperB.

- `smoking guns' for their models.
- •Measurements not yet made are denoted by X.

With 75ab<sup>-1</sup> we can
Reach above a TeV with B→ τν
See B→Kνν

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### **Accelerator Aspects**

How can we obtain a data sample of 75ab<sup>-1</sup>?

## Target Integrated Luminosity

- Why at least 75ab<sup>-1</sup> of data?
  - Many of these new physics searches become systematically or theoretically limited.
    - e.g. time dependent asymmetry measurements with  $b \rightarrow s$  penguin decays).
  - This data sample represents almost two orders of magnitude larger sample than current experiments.
    - The current B-factories have over 1ab<sup>-1</sup> (combined) on disk/tape.
    - Will record a total of ~1.5ab<sup>-1</sup>.
  - Ensures that if new physics is found (e.g. in LFV) that one can start to perform rudimentary measurements of such phenomena.
    - 10ab<sup>-1</sup> of data is sufficient to start to constrain models of LFV in  $\tau$  decays.
    - Need more than this to ensure discovery.
  - Will be able to start measuring parameters in V<sub>SCKM</sub> (if SUSY exists), or constrain Multi TeV energy level NP in your favourite scenario.
    - Strong constraints on NP that complement the LHC direct searches!
  - Will be able to test for light Higgs/dark matter particles and lepton universality by running at the Y(3S) resonance [hundreds of fb<sup>-1</sup> from a few months running].

## How to get increased $\mathcal{L}$



- Option 1: Brute Force.
  - Increase beam current.
  - Decrease  $\beta^*_{y}$ .
  - Increase beam-beam effect  $\xi$  (reduce bunch length).

(Hard – but possible – to do all of this efficiently)

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## How to get increased $\mathcal{L}$



- Have a 15mrad crossing angle of beams.
- Focus beams at IP (small  $\beta^*$ ).
- Retain longer bunch lengths.
- Rotate colliding bunches so no geometric loss at IP.
  - Align the focussed parts of bunches that cross each other at the IP. Call this "Crab Crossing/Waist".

## Large crossing angle, small x-size



### Crab waist tests at $DA\Phi NE$



#### PARAMETERS

#### J.Seeman @MiniMac

LER/HER	Unit	June 2008	Jan. 2009	March 2009	LNF site	
E+/E-	GeV	4/7	4/7	4/7	4/7	
L	cm2 s1		1x10 <sup>10</sup>	1x10 <sup>H</sup>	1x10 <sup>16</sup>	
17/1	Amp	1.85 /1.85	2.00/2.00	2.80/2.80	2.70/2.70	
Npart	x10 <sup>10</sup>	5.55 /5.55	6/6	4.37/4.37	4.53/4.53	
Noun		1250	1250	2400	1740	
burch	mA	1.48	1.6	1.17	1.6	
6/2	mrad	25	30	30	30	
B.*	mm	35/20	35/20	35/20	35/20	
B,*	mm	0.22 /0.39	0.21 /0.37	0.21 /0.37	0.21 /0.37	
ε,	nm	2.8/1.6	2.8/1.6	2.8/1.6	2.8/1.6	
е,	pm	7/4	714	714	7/4	
σ,	μm	9.9/5.7	9.9/5.7	9.9/5.7	9.9/5.7	
α,	nm	39/89	38/38	38/38	38/38	
a,	mm.	515	5/5	5/5	5/5	
د	X tune shift	0.007/0.002	0.005/0.0017	0.004/0.0013	0.004/0.0013	
e,	Y tune shift	0.14 /0.14	0.125/0.126	0.091/0.092	0.094/0.095	
RF stations	LERIHER	5/6	5/6	5/8	6/9	
RF wall plug power	MW	16.2	18	25.5	30.	
Circumference	m	1800	1800	1800	1400	

### SITES : Tor Vergata.....



Perugia June 15,2009

Marcello A. Giorgi

### LNF option


... "Mini-MAC now feels secure in enthusiastically encouraging the SuperB design team to proceed to the TDR phase, with confidence that the design parameters are achievable"

Machine is possible! The 2 rings can be built largely with the components of PEPII:

Magnets and RF stations.

## **Detector Design**



# Tracking



BaBar DCH Design

- Adequate performance.
- Needs to be replaced as the existing detector is aging.



## All Pixel SVT Concept

- Use INMAPS chips for a 5 layer all pixel vertex detector.
  - Adapt well understood leading STFC funded design to use with SuperB.
  - Common infrastructure for sub -system.
  - Physics studies required to understand performance (in progress) as part of detector optimisation.
  - UK has world leading expertise in this area.
  - Building on expertise and developments from SPiDeR and CALICE, LCFI ...
  - Concept well received by SuperB.





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## Particle ID

- Detector of Internally Reflected Cherenkov light (DIRC) works extremely well.
- Aim to reuse this principle with state of the art readout.





Can benefit from reducing the volume of water between the end of the quartz bars and the photodetectors (PMTs) at SuperB.

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## **Calorimeter Barrel**

Calorimeter Barrel is more than sufficient for our needs.
 Fast enough signal output for the



# **Calorimeter End-Cap**

- BaBar End-Cap doesn't have a fine enough granularity for rates at SuperB.
  - Need a finer segmentation.
  - Similar total X<sub>0</sub>.
  - Faster readout electronics.
  - Several candidate materials for End-Cap replacement.
    - LYSO is baseline
      - expensive at the moment (~ \$40/cc).
      - Aim for \$15/cc.
  - Need to integrate into the existing Barrel, and optimise segmentation.
  - R&D underway toward a LYSO Calorimeter test-beam in ~2009.





## **Instrumented Flux Return**

- BaBar has 5 radiation lengths of material for  $\mu$  identification in the • flux return.
  - This is not optimal. \_
  - SuperB will have more iron.
- The segmentation of active regions of the flux return will remain the • same as BaBar (3.7cm pitch).
- 7-8 layers of MINOS style scintillator bars. •





ATTENUATION LENGTH MEASUREMENTS FOR 5 CASES

CURVE IS FIT TO **YPONENTIA** 

CASES 48 5

200

45

50

250

### **Detector Simulation**

- Simulation:
  - FastSim (validated on using geometry for BaBar)
    - Reproduces BaBar resolutions etc.
    - Change to SuperB geometry and boost for development of benchmark studies.
    - Then move to GEANT 4 for more detailed work.
  - GEANT 4 model of SuperB shown.
  - Using BaBar framework.
  - Draw on a decade of analysis experience from BaBar and Belle to optimize an already good design.



## Current status

### Timeline of the project



Prior to 2005, there was no clear way to achieve an interesting data sample on an interesting timescale ( $\mathcal{L} < 10^{36} \text{ cm}^{-2}\text{s}^{-1}$ ).

Then there was a revelation: The crabbed waist and inspiration from the ILC.

Working toward a coherent description of what we want to build and why: White paper end of 2009 Technical Design Reports end of 2010.

Expect a funding decision from host country by the end of this year.

5 years of nominal data taking will give 75ab<sup>-1</sup> of data.

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#### **Global situation**

- Global community is working on the SuperB TDR effort: Accelerator, Detector, Physics. Global community working
- R&D funding from:
  - Italy
  - France
  - Canada

Global community working toward the realization of a Super Flavour Factory.

Two Concepts: SuperB (this seminar) and Belle-II.

- Other R&D efforts at various levels from:
  - Israel, Poland, Norway, Russia, Spain, UK, Ukraine, US.
- Other countries are watching the development of the process.
- Please contact me if you want more information or want to join the effort: a.j.bevan@qmul.ac.uk.

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#### What about Belle-II?

- Similar concept:
  - Adiabatic upgrades from KEKB through to a  $\sim 0.8 \times 10^{36}$  machine.
    - Funding situation was equivalent to SuperB.
    - Timeline is start data taking in 2013 (low luminosity).
    - Incremental upgrades to reach the ultimate lumi.
    - Target data sample: 50ab<sup>-1</sup>.
  - Some differences between SuperB and Belle-II by ~2020:

Experiment:	SuperB	Belle-II
I <sub>LER</sub>	4 GeV	tbd
I <sub>HER</sub>	7 GeV	tbd
٤ <sub>x</sub>	2.8 / 1.6 nm	tbd
ε <sub>y</sub> Γ	7 / 4 pm	tbd
Ĺ	75ab <sup>-1</sup>	50ab⁻¹
e <sup>-</sup> Polarisation	80%	none
run at ψ(3770)	yes	no

N.B. Some parameters for the experiments may change. The Belle-II accelerator concept is in the process of being re-worked from a high current to a low emmitance (Italian) one, so the total cost of both projects will be the about the same.

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#### What about Belle-II?

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Experiment:	SuperB	Belle-II
Ι <sub>LER</sub> Ι <sub>HER</sub> ε <sub>x</sub>	4 GeV 7 GeV 2.8 / 1.6 nm	Polarisation increases potential of T physics studies.
ε <sub>y</sub>	7 / 4 pm	ψ(3770) increases charm/CPV
L	75ab <sup>-1</sup>	/Mixing study potential.
e⁻ Polarisation	80%	none
run at ψ(3770)	yes	no

N.B. Some parameters for the experiments may change. The Belle-II accelerator concept is in the process of being re-worked from a high current to a low emmitance (Italian) one, so the total cost of both projects will be the about the same.

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Hindsight always gives us 20:20 vision.

Until we have understood new physics, we are left trying to piece together the jigsaw puzzle of a high energy world where the possibilities are limited only by (a theorists) imagination.

- Want to elucidate new physics in as many ways as possible. Currently we
  - don't know the fine detail of NP.
  - don't know the relevant NP energy scale (yet).
    - The LHC may, or may not elucidate this issue.
  - don't know if the NP flavour sector is trivial or complicated:
    - Prior experience suggests it will be complicated.
  - But we do know that there are many models: 2HDM (type-n), MSSM, NMSSM, ...
    - Many *assume* flavour couplings are zero.

- The LHC won't be able to solve the SUSY flavour problem.
  - LHCb may help in a few specific channels: e.g. K\*II,  $B_{\rm S}$  decays.
  - The GPDs may help with some ultra-rare B decays.
  - Some NP sensitive observables are accessible through studies at dedicated flavour experiments.
  - A large number of observables are only measureable competitively at a Super Flavour Factory.
    - Need this to unravel the nature of new physics.

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 A subset of interesting light meson flavour-physics circa 2015: Predictions from Koppenberg: '09.

#### Summary

		Improved	Limited	Interesting	
Mo	de	by 2014	by theory	In 2014?	Where
$B_d$	$\rightarrow \mu\mu$	Yes	No	Yes	pp
→ B -	$\rightarrow  au  u$	Hardly	No	Yes	$\Upsilon(4S)$
🔶 b –	+ $s\gamma$ polarisation	Yes	No	Yes	Both
<b>B</b> -	$\rightarrow \ell \ell K^*$	Yes	No	Yes	pp
→ B -	$\rightarrow \nu \nu K$	Hardly	No	Yes	$\Upsilon(4S)$
🔶 CP	$(b  ightarrow s \gamma  ightarrow b  ightarrow d \gamma)$	No	No	Yes	$\Upsilon(4S)$
Exe	clusive $b  ightarrow d$	Yes	No	Yes	Both
BF	$(b  ightarrow s \gamma)$	Hardly	Yes(?)	Unlikely	$\Upsilon(4S)$
$B_s$	$\rightarrow \mu \mu$	Yes	Yes	No	
	decays	Yes	Maybe	Unclear	Both
K	$\rightarrow \pi \nu \nu$	Yes	No	Yes	Dedic.
erial Coll	ege				37

**Caveats I:** 

Focuses on what can be done now.

SuperB would open the doors to many new interesting modes.

Most interesting measurements are possible at SuperB.

Read in conjunction with the next slide.

March 2008

 A subset of interesting light meson flavour-physics circa 2015: Predictions from Koppenberg: '09.

#### Summary

	Improved	Limited	Interesting	
Mode	by 2014	by theory	In 2014?	Where
$B_d \rightarrow \mu \mu$	Yes	No	Yes	pp
$B \rightarrow \tau \nu$	Hardly	No	Yes	$\Upsilon(4S)$
$b  ightarrow s \gamma$ polarisation	Yes	No	Yes	Both
$B  ightarrow \ell \ell K^*$	Yes	No	Yes	pp
$B \rightarrow \nu \nu K$	Hardly	No	Yes	$\Upsilon(4S)$
$ \qquad \qquad$	No	No	Yes	$\Upsilon(4S)$
Exclusive $b \rightarrow d$	Yes	No	Yes	Both
$BF(b o s\gamma)$	Hardly	Yes(?)	Unlikely	$\Upsilon(4S)$
$B_s  ightarrow \mu \mu$	Yes	Yes	No	
D decays	Yes	Maybe	Unclear	Both
$K \rightarrow \pi \nu \nu$	Yes	No	Yes	Dedic.
erial College				2
don P. Koppenburg	Rare Decays-	13/07/09 - P	PAP community	review - p

#### Caveats II:

Inclusive measurements are theoretically clean: so they are more interesting to make.

You need a clean environment (e<sup>+</sup>e<sup>-</sup>) to do inclusive measurements: i.e. SuperB

Quantum correlations at the  $\psi(3770)$  open up an equivalent set of measurements at charm threshold.

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- The rest of the golden matrix includes:
  - more T Lepton Flavour Violation studies.
  - Y decay studies:
    - Light Higgs
    - Dark Matter

	$H^+$ high tan $\beta$		Non-MFV	NP Z-penguins	Right-handed currents	LTH	SUSY
$\mathcal{B}(B \to X_s \gamma)$		L	M		M		
$\mathcal{A}_{CP}(B \to X_s \gamma)$			L		M		
$\mathcal{B}(B \to \tau \nu)$	L-CKM						
$\mathcal{B}(B \to X_s \ell \ell)$			M	M	M		
$\mathcal{B}(B \to K \nu \overline{\nu})$			M	L			
$S_{K_S \pi^0 \gamma}$					L		
The angle $\beta$ ( $\Delta S$ )			L-CKM		L		
$\tau \rightarrow \mu \gamma$							L
$\tau \rightarrow \mu \mu \mu$						L	

- Lepton number conservation
- Not discussed here. Probe Dark Sector: 'Dark Forces'
  - Learn about simple, and complex models of Dark Matter through meson decays.
  - http://www-conf.slac.stanford.edu/darkforces2009/

- Many questions will remain after the anticipated discoveries from the LHC.
  - What is the nature of flavour couplings beyond the Standard Model?
    - What can this tell us about any (un-)observed new particles?
  - Are there additional contributions that help resolve the Matter-antimatter asymmetry fine tuning problem?
  - Charged Lepton Flavour Violation: is it SM ~10<sup>-54</sup>, or significantly enhanced?
  - Is there a charged Higgs?
  - What is the structure of Dark Matter?
- SuperB can shed some light in all these areas.

```
March 2008
```



#### All we need to do is build it!

#### New effort is welcome!

#### http://web.infn.it/superb/index.php/home



## **Extra Material**

### **Precision CKM**

- CKM is a 36 year old anzatz.
- Works at the 10% level.
- No underlying physical insight.
- Small new physics contributions not ruled out (% level).

Precision CKM from SuperB will open up more new physics search opportunities: e.g.  $K \rightarrow \pi \nu \nu$ :



K+ decay has a similar error budget.



## Particle Physics Landscape circa 2015



#### **Dark Forces**



See the recent workshop <u>http://www-conf.slac.stanford.edu/darkforces2009/</u> Summarised by Mat Graham at the October 2009 SLAC SuperB meeting Arkani-Hamed, Finkbeinder, Slatyer,



#### **Dark Forces**



See the recent workshop <a href="http://www-conf.slac.stanford.edu/darkforces2009/">http://www-conf.slac.stanford.edu/darkforces2009/</a> Summarised by Mat Graham at the October 2009 SLAC SuperB meeting

In addition to the vector 'portal' with the kinetic coupling, there should be a Higgs coupling term:

•B→K\*4I is an interesting channel to search for this.



Adrian Bevan http://www.pi.infn.it/SuperB/

## New Physics in ∆F=2 Transitions

•  $\Delta$ F=2 transitions in  $B^0_{d,s}\overline{B}^0_{d,s}$  systems are box diagrams (mixing or FCNC).



$$q = d, s$$

• New physics (NP) can contribute with an amplitude ratio  $C_q$  and phase  $\phi_{q_1}$ 

$$C_{q}e^{i\phi_{q}} = \frac{\left\langle B_{q}^{0} \mid H_{SM+NP} \mid \overline{B}_{q}^{0} \right\rangle}{\left\langle B_{q}^{0} \mid H_{SM} \mid \overline{B}_{q}^{0} \right\rangle}$$

•  $C_q=1$ , and  $\phi_q=0$  for the Standard Model (SM).

## New Physics in ∆F=2 Transitions

- Existing measurements already constrain NP in B<sub>d</sub> mixing (See later for B<sub>S</sub>).
- SuperB will significantly improve this constraint.



# **Minimal Flavour Violation**

- Suppose that there are no new physics flavour couplings (MFV).
  - CP violation comes from the known SM Yukawa couplings.
  - The top quark contribution dominates the SM.
  - NP contribution in  $\Delta B=2$  transitions is:



ADD, RS.

• What is the energy scale that we are sensitive to?

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## Minimal Flavour Violation

- Sensitive to new physics contributions with  $\Lambda$  up to 14 TeV (=  $6\Lambda_0$ ).
- For loop mediated NP contributions the constraint can be weakened so that  $\Lambda \sim 700$ GeV.
- Don't require that the EWSB scale match  $\Lambda$ .

## Aside: MFV & B<sub>S</sub>?

- Recent preprint from UT Fit claims evidence for new physics in B<sub>S</sub> decays.
  - Test for NP via:

$$C_{s}e^{i\phi_{s}} = \frac{\left\langle B_{s}^{0} \mid H_{SM+NP} \mid \overline{B}_{s}^{0} \right\rangle}{\left\langle B_{s}^{0} \mid H_{SM} \mid \overline{B}_{s}^{0} \right\rangle}$$

- Using B<sub>S</sub> mixing, A<sub>SL</sub>, lifetime and tagged J/ $\psi \phi$  results ( $\Delta \Gamma$  vs  $\beta_S$ ) from CDF and D0.

$$\beta_{s} = 0.018 \pm 0.001 \text{ (SM)}$$
$$= \arg\left(\frac{-V_{ts}V_{tb}^{*}}{V_{cs}V_{cb}^{*}}\right)$$

Adrian Bevan http://www.pi.infn.it/SuperB/ e.g. see arXiv:0803.0659 69

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  - Test for NP via:



Observable	68% Prob.	95% Prob.
$\phi_{B_s}[^\circ]$	$-19.9 \pm 5.6$	[-30.45, -9.29]
	$-68.2 \pm 4.9$	[-78.45, -58.2]
$C_{B_s}$	$1.07\pm0.29$	[0.62, 1.93]
$\phi_s^{\mathrm{NP}}[^\circ]$	$-51 \pm 11$	[-69, -27]
	$-79\pm3$	[-84,-71]
$A_s^{ m NP}/A_s^{ m SM}$	$0.73\pm0.35$	[0.24, 1.38]
	$1.87\pm0.06$	[1.50, 2.47]
$\mathrm{Im} \; A_s^{\mathrm{NP}} / A_s^{\mathrm{SM}}$	-0.74 $\pm$ 0.26	[-1.54, -0.30]
${\rm Re}~A^{\rm NP}_s/A^{\rm SM}_s$	$-0.13\pm0.31$	[-0.61, 0.78]
	-1.82 $\pm$ 0.28	[-2.68, -1.36]
$A_{ m SL}^s  imes 10^2$	$-0.34 \pm 0.21$	[-0.75, 0.03]
$A_{\rm SL}^{\mu\mu}  imes 10^3$	$-2.1 \pm 1.0$	[-4.7, -0.3]
$\Delta \Gamma_s / \Gamma_s$	$0.105\pm0.049$	[0.02, 0.20]
	$-0.098 \pm 0.044$	[-0.19, -0.02]

$$\beta_s = 0.0409 \pm 0.0038$$

Eagerly awaiting a final result from CDF and D0: AND results from LHCb!

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# SUSY CKM

- The SM encodes quark mixing in the CKM matrix,  $\nu$  mixing with the MSW matrix .... so
- SUSY encodes squark mixing in a Super CKM equivalent of the CKM matrix: V<sub>SCKM</sub>.

Let us now consider a MSSM with generic soft SUSY-breaking terms, but dominant gluino contributions only  $\binom{d}{s}^{d}{y}$ 



- Have couplings for LL, LR, RL, RR interactions.
- LHC probes the High Energy Frontier.
  - Measures the diagonal elements of  $V_{SCKM}$ .
- SuperB probes the Luminosity Frontier.
  - Measures the off-diagonal elements  $V_{\text{SCKM}}$ .
# SUSY CKM

- Couplings are  $(\delta_{ij}^{q})_{AB}$  L. Silvestrini (SuperB IV) where A,B=L,R, and i,j are squark generations.
- e.g. Constrain parameters

in  $V_{\text{SCKM}}$  using:

•  $\mathcal{B}(B \rightarrow X_s \gamma)$  [green] •  $\mathcal{B}(B \rightarrow X_s |^+l^-)$  [cyan] •  $A_{CP}(B \rightarrow X_s \gamma)$  [magenta] • Combined [blue]

SuperB probes new physics in SUSY larger than 20TeV (and up to 300TeV in some scenarios)



With current data, the whole range shown is allowed!



Adrian Bevan http://www.pi.infn.it/SuperB/

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# D<sup>0</sup> mixing

- Recent measurements from **BaBar and Belle** demonstrated B factory capabilities in charm physics
- Possibility to measure CP violation in the charm sector •





#### · · · · · ·



, Projected Sensitivity									
Exp't / 1 o	y <sub>Ф</sub> (10-3)	y' (10-3)	x'2 (10-4)	cosõ					
B-factories (2ab <sup>-1</sup> )	2-3	2-3	1-2	-					
SuperB (50 ab <sup>-1</sup> )	0.5	0.7	0.3	-					
LHCb (10 fb-1) Only B->D*	?	0.7	0.7	-					
LHCb (100 fb-1) Prompt D*	?	?	?	-					
CLEO-c (750 pb-1)	10		2-3	0.1-0.2					
BESIII (20 fb-1)	4		0.5-1	0.05					
Super8 - 4 GeV (0.2 ab <sup>-1</sup> )	1-2	-	<0.2	<0.05					

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IIIIp.// WWWW.pl.IIIII.IU OuporD/

# Searching for a Light Higgs or Dark Matter Candidates

LEP data do not exclude the possibility!

For more details see the talks of McElrath and Sanchis at the SuperB retreat in Valencia Jan '08: http://ific.uv.es/superb/

# Searching for a Light Higgs

- Many NP scenarios have a possible light Higgs Boson (e.g. 2HDM).
- Can use  $Y(nS) \rightarrow I^+I^-$  to search for this.
  - Contribution from A<sup>0</sup> would break lepton universality



M. A. Sanchis-Lozano, hep-ph/0510374, Int. J. Mod. Phys. A19 (2004) 2183

Can expect hundreds of fb<sup>-1</sup> recorded at the Y(3S) in SuperB

• NMMSM Model with 7 Higgs Bosons

#### Physical Higgs bosons: (seven)

- 2 neutral CP-odd Higgs bosons  $(A_{1,2})$ 3 neutral CP-even Higgs bosons  $(H_{1,2,3})$ 2 charged Higgs bosons  $(H^{\pm})$
- $A_1$  could be a light DM candidate.

Possible NMMSM Scenario

 $A_1 \sim 10 \text{ GeV}$  $H_1 \sim 100 \text{ GeV}$  (SM-like) Others ~300 GeV (almost degenerate)

Gunion, Hooper, McElrath [hep-ph:0509024] McElrath [hep-ph/0506151], [arXiv:0712.0016]

### **Searching for Dark Matter**



Adrian Bevan http://www.pi.infn.it/SuperB/

### $\tau \text{ Decays}$

### $\tau \rightarrow \mu \gamma$ / 3leptons

 Comparison of μ→eγ and τ→μγ rates can distinguish between NP scenarios.



- Can depend on the value of  $\theta_{13}$ .
- Best search capability for LFV in τ→3leptons of any experiment.



# CP and CPT Violation

- <sup>Studies</sup> starting
- SM decays of the  $\tau$  have only a single amplitude so any CP violation signal is an unambiguous sign of NP. Can have NP contributions from a UP.

  - Can have NP contributions from a H<sup>±</sup> in many modes, and largely experimentally un-explored.

e.g. see Datta et al., hep-ph/0610162

- CPT Violation.
  - Expect to be able to measure  $\frac{\tau_{\tau_{-}} \tau_{\tau_{+}}}{\tau_{\tau_{-}} + \tau_{\tau_{+}}}$  at the level of 10<sup>-4</sup> (statistical).
  - Current bound is  $(0.12 \pm 0.32)$ %.

Nucl. Phys. Proc. Suppl. 144 105 (2005)

 Polarisation of e<sup>+</sup>e<sup>-</sup> beams benefits the search for CP and CPT violation in  $\tau$  decay and the  $\tau$ anomalous magnetic moment. e.g. PRD 51 3172 (1995); arXive :0707.2496 [hep-ph]

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# **Detector Design**

#### Requirements

- The B-factory detectors work extremely well.
  - Design of a SuperB detector, essentially means a refinement of the existing detectors.
- SuperB environment will have a higher rate.
  - Some existing detector parts are reusable.
    - Csl Calorimeter barrel.
    - DIRC quartz bars from BaBar. These 3m long bars are required for the particle identification system.
    - Superconducting Solenoid Magnet: creates a 2T magnetic field.
  - Some existing detector parts need to be replaced to cope with the expected rates.
    - Central tracking inside the particle ID system.
    - End Cap of the calorimeter.
    - Instrumented Flux Return ( $\mu$ , K<sup>0</sup><sub>L</sub> detector).
    - Readout electronics.
  - Makes sense to optimise reuse in order to limit the cost of the project.

# DAQ

#### Modelled on the BaBar Data Acquisition system.

As is the norm with modern

experiments, will need tens



### Timescale

- Overall schedule dominated by:
  - Site construction.
  - PEP-II/Babar disassembly, transport, and reassembly.
- Possible to reach the commissioning phase after 5 years from T0.
- Physics from circa 2015?



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http://www.pi.inf

Figure 5-1. Overall schedule for the construction of the SuperB project.

### Accelerator and site costs

		EDIA	Labor	<i>M</i> \&S	Rep.Val.
WBS	ltem	mm	mm	kEuro	kEuro
1	Accelerator	5429	3497	191166	126330
1.1	Project management	2112	96	1800	0
1.2	Magnet and support system	666	1199	28965	25380
1.3	Vacuum system	620	520	27600	14200
1.4	RF system	272	304	22300	60000
1.5	Interaction region	370	478	10950	0
1.6	Controls, Diagnostics, Feedback	963	648	12951	8750
1.7	Injection and transport systems	426	252	86600	18000
		EDIA	Labor	<i>M</i> \&S	Rep.Val.
WBS	Item	тт	тт	kEuro	kEuro
2.0	Site	1424	1660	105700	0
2.1	Site Utilities	820	1040	31700	0

Note: site cost estimate not as detailed as other estimates

**Tunnel and Support Buildings** 

Funds needed to build experiment

620

604

Replacement value of parts that we can re-use.

0

74000

2.2

#### **Detector cost**

WBS	ltem	EDIA mm	Labor mm	M∖&S kEuro	Rep.Val. kEuro
1 1	SuperB detector	3391	1873	40747	46471
1.0	Interaction region	10	4	210	0
1.1	Tracker (SVT + L0 MAPS)	248	348	5615	0
1.1.1	SVT	142	317	4380	0
1.1.2	L0 Striplet option	23	33	324	0
1.1.3	L0 MAPS option	106	32	1235	0
1.2	DCH	113	104	2862	0
1.3	PID (DIRC Pixilated PMTs + TOF)	110	222	7953	6728
1.3.1	<b>DIRC barrel - Pixilated PMTs</b>	78	152	4527	6728
1.3.1	DIRC barrel - Focusing DIRC	92	179	6959	6728
1.3.2	Forward TOF	32	70	3426	0
1.4	EMC	136	222	10095	30120
1.4.1	Barrel EMC	20	5	171	30120
1.4.2	Forward EMC	73	152	6828	0
1.4.3	Backward EMC	42	65	3096	0
1.5	IFR (scintillator)	56	54	1268	0
1.6	Magnet	87	47	1545	9623
1.7	Electronics	286	213	5565	0
1.8	Online computing	1272	34	1624	0
1.9	Installation and integration	353	624	3830	0
1.A	Project Management	720	0	180	0

Note: options in italics are not summed. We chose to sum the options we considered most likely/necessary.

#### Total = 338M Euro.

= 510M Euro (counting the cost of re-used parts).
⇒ 1/3 of the cost of the project can be saved by re-using parts of BaBar and PEP-II.