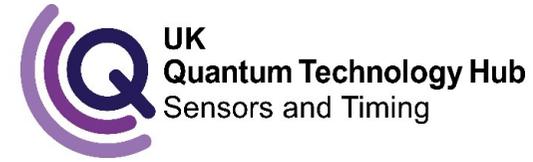




UNIVERSITY OF  
BIRMINGHAM



# Atom Interferometry for Fundamental Physics

**Dr. Samuel Lellouch**

**Particle Physics Seminar, March 8<sup>th</sup> 2023**

# Outline

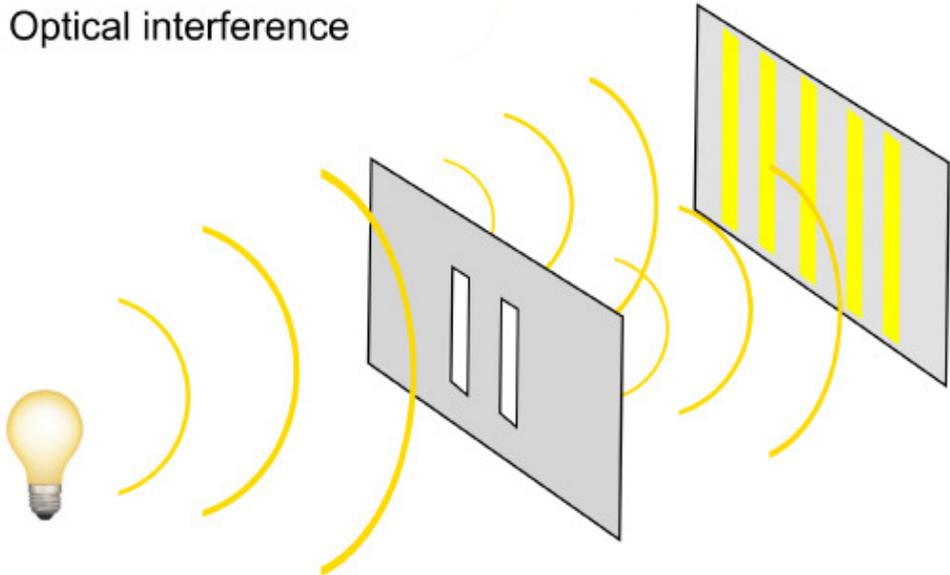
- I. Basics of atom interferometry
- II. Practical applications in quantum sensing
- III. Fundamental physics applications: the AION project

# **PART I.**

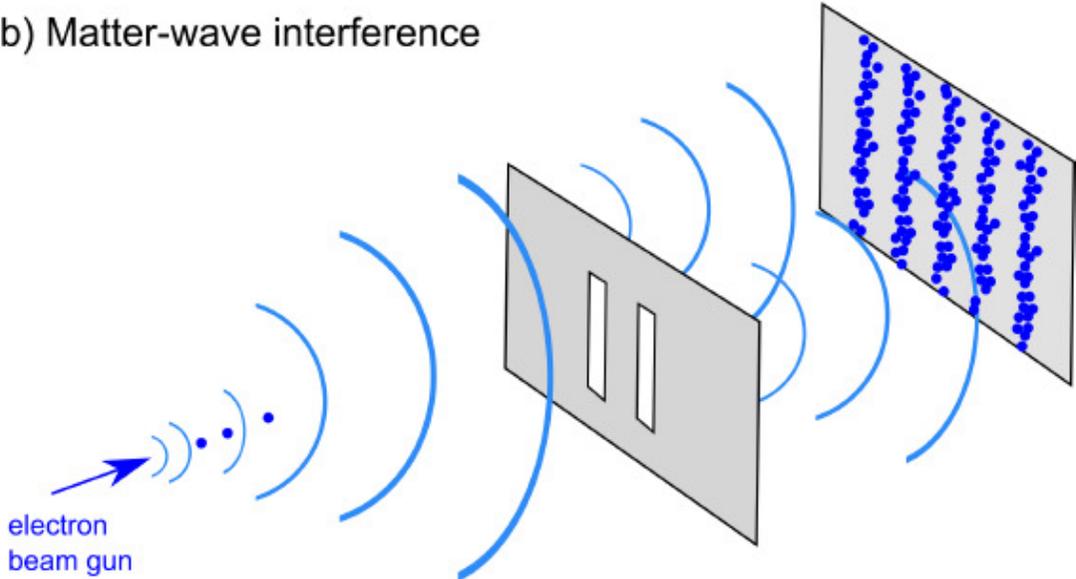
## **BASICS OF ATOM INTERFEROMETRY**

# Interference of matter-waves

a) Optical interference

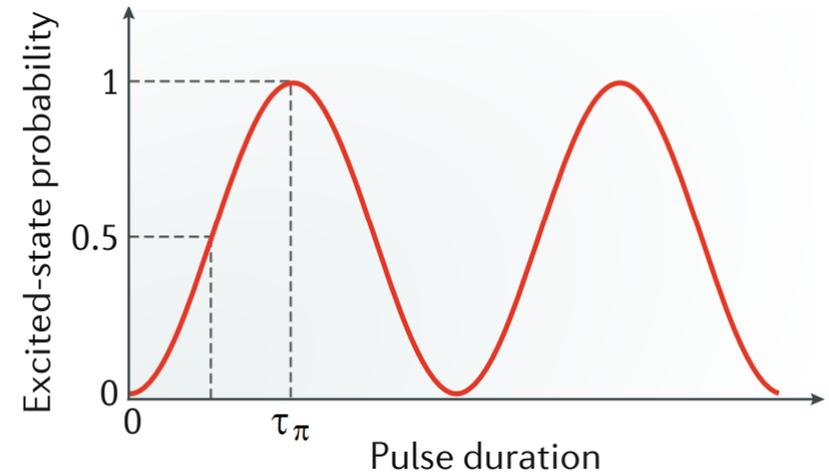
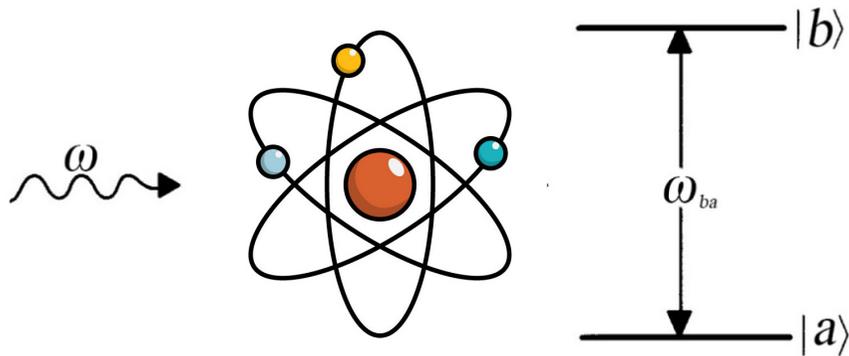


b) Matter-wave interference



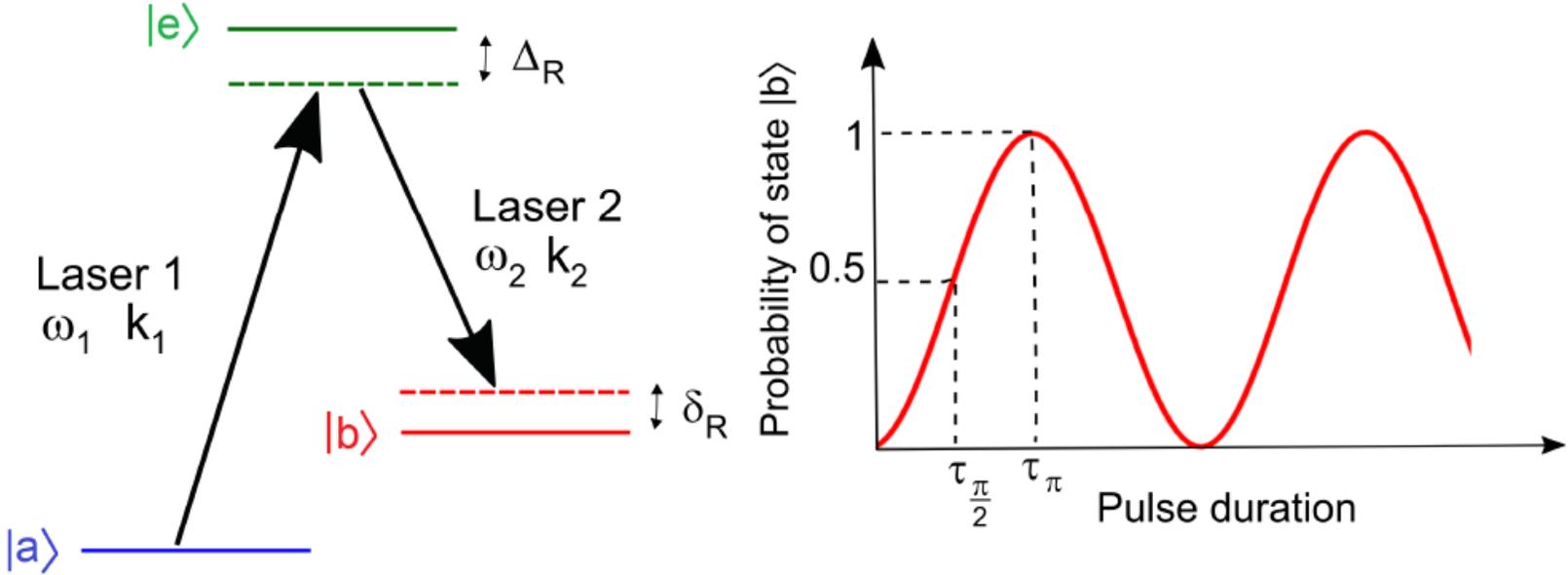
# Manipulating atoms with light pulses

## The two-level atom: Rabi oscillations

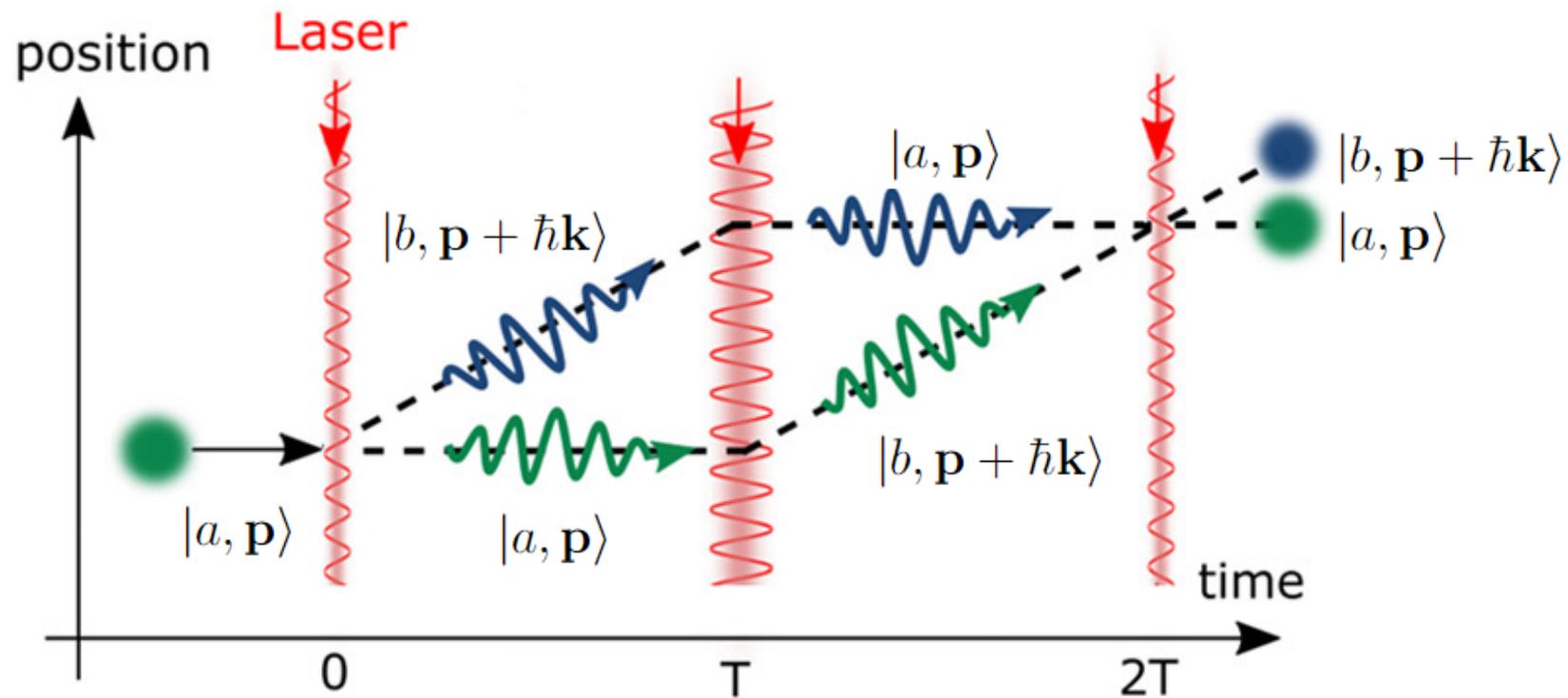


# Manipulating atoms with light pulses

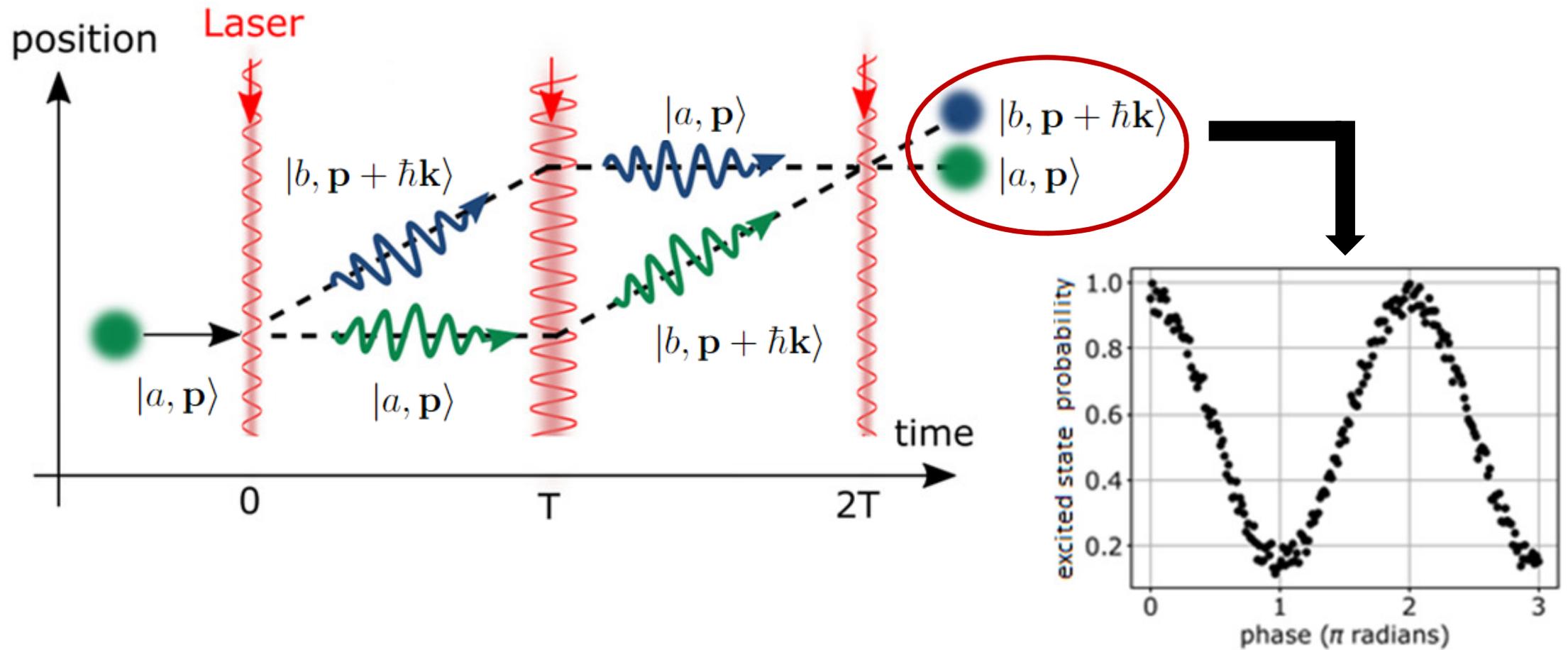
## Stimulated Raman transitions



# Atom interferometry

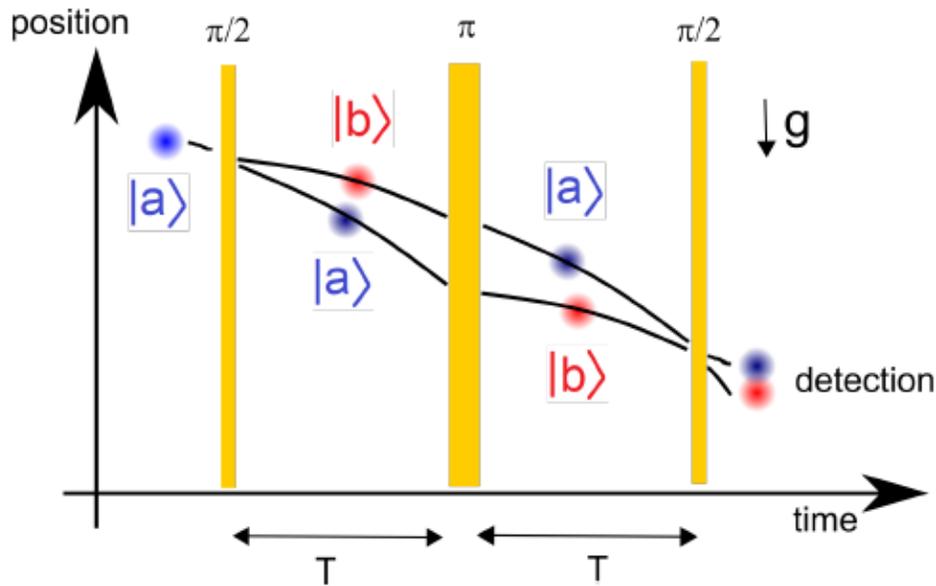


# Atom interferometry

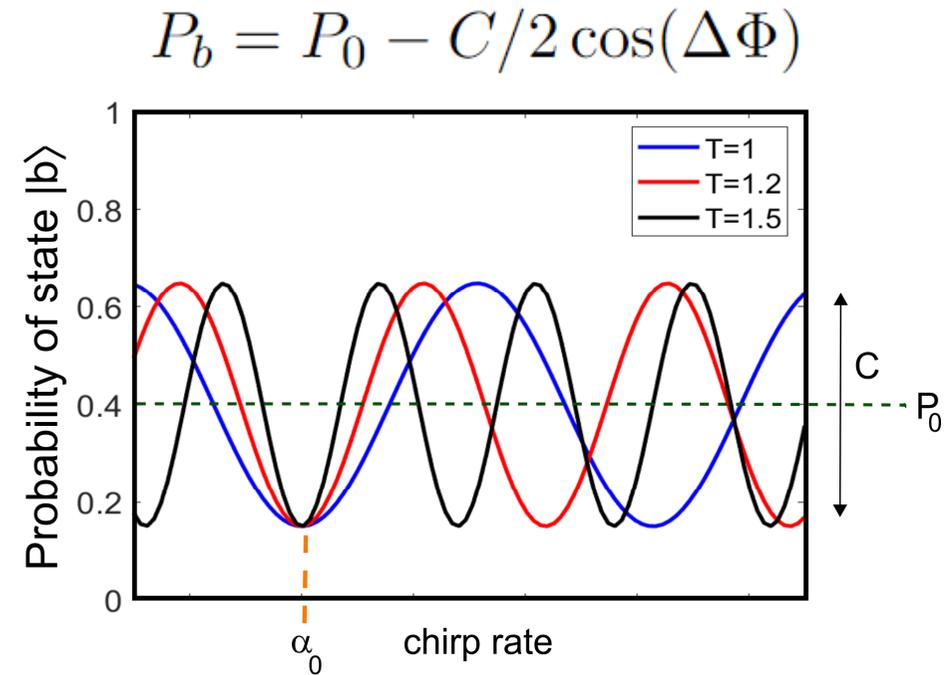


$$P_b = P_0 - C/2 \cos(\Delta\Phi)$$

# Example: performing a gravity measurement



In free fall under gravity:  $\Delta\Phi = k_{\text{eff}} \cdot g T^2$



In practice, we chirp the laser frequency linearly with time,  $\omega \rightarrow \omega + \alpha t$ , and we determine the chirp rate  $\alpha_0$  which cancels out the gravity phase shift for any  $T$ :  $g = 2\pi\alpha_0/k_{\text{eff}}$

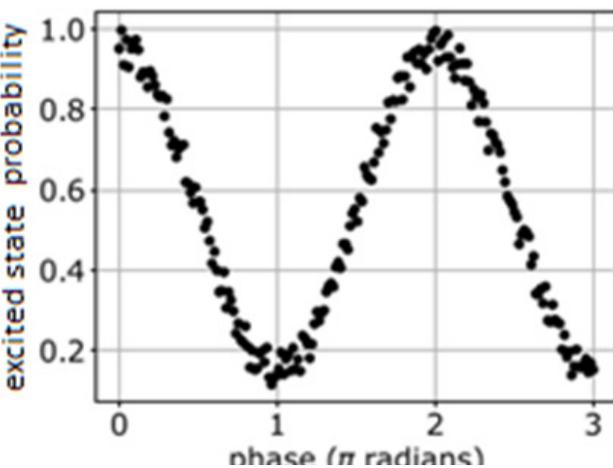
# Advantages of atom interferometry

$$\Delta\Phi = k_{\text{eff}} \cdot g T^2$$

**Accuracy:** Arises from the large scale-factor of the interferometer,  $k_{\text{eff}} T^2$

**Long-term stability:** The scale factor is immutable and controlled with high precision

**Sensitivity:** In the absence of any noise source, the sensitivity is limited by the quantum projection noise



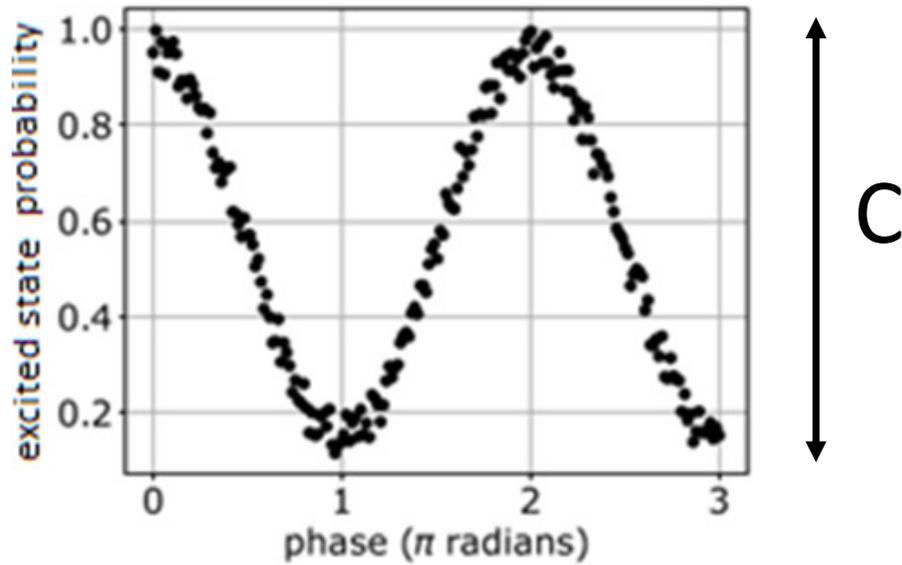
$$\sigma_g = \frac{1}{C k_{\text{eff}} T^2 \sqrt{N}}$$

Enhance contrast

Increase momentum  
separation

Increase interrogation time

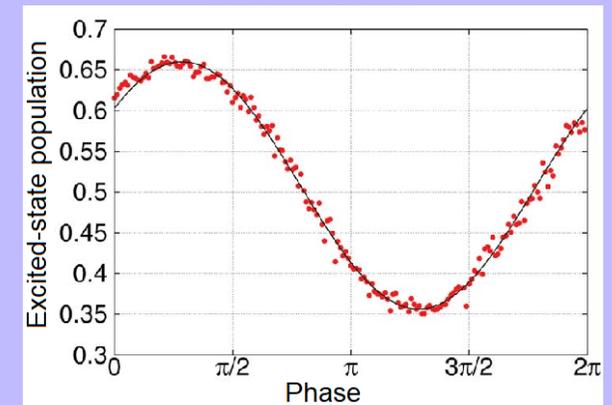
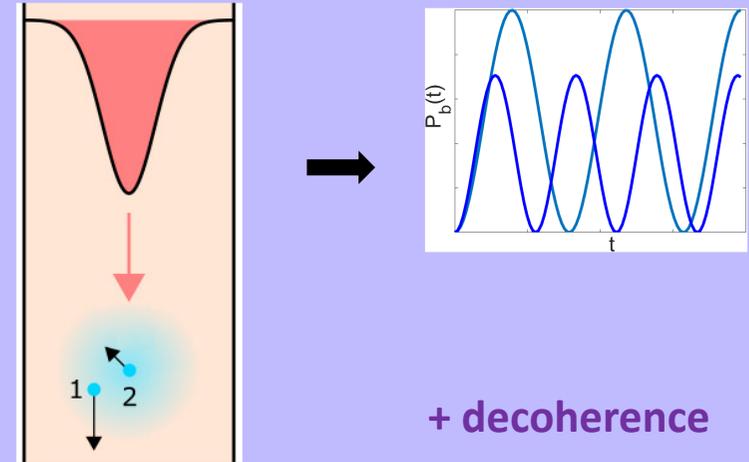
# Interferometer contrast



$$P_b = P_0 - C/2 \cos(\Delta\Phi)$$

## Contrast losses

### Cloud inhomogeneities

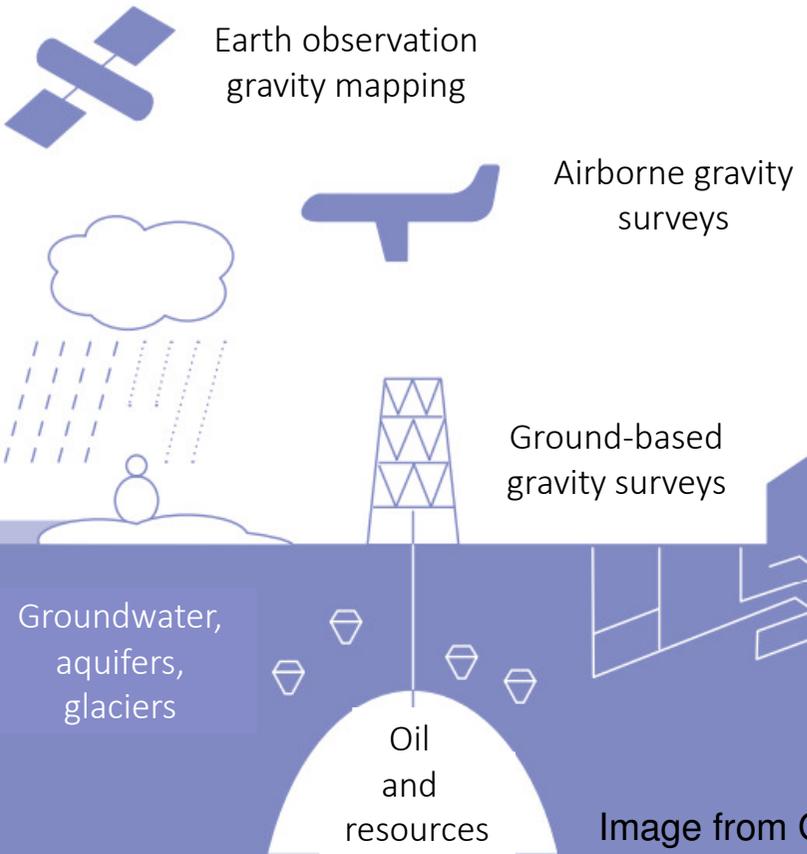


## **PART II.**

# **PRACTICAL APPLICATIONS IN QUANTUM SENSING**

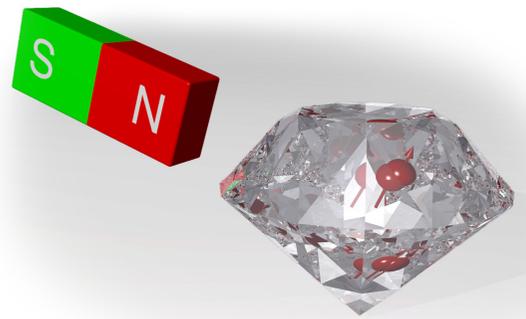
# Quantum sensing technologies

## Gravimetry



Key infrastructure

## Magnetometry



## Inertial sensing

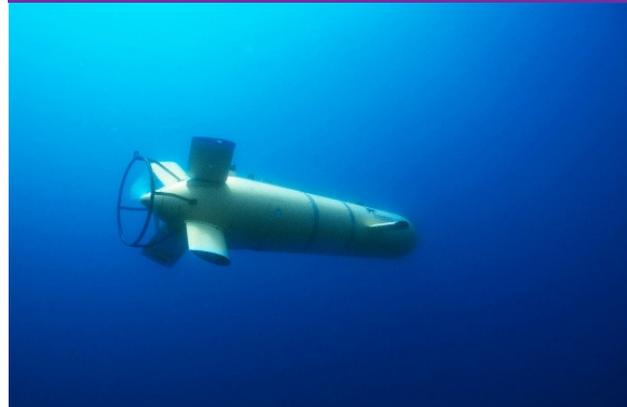
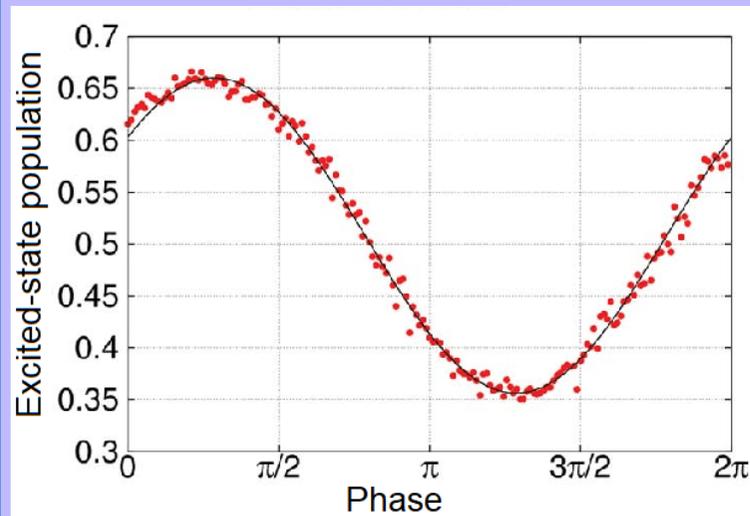


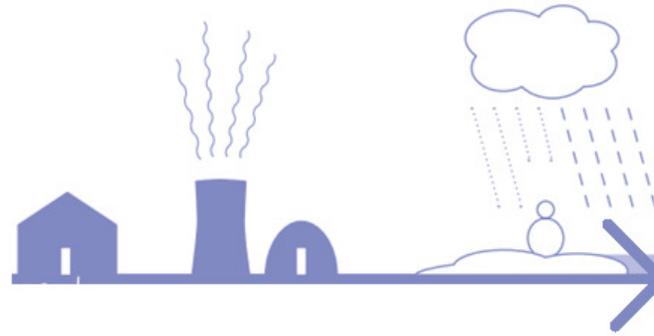
Image from Quantum technologies: Blackett review, 2016

# From the lab to the field

## In lab conditions

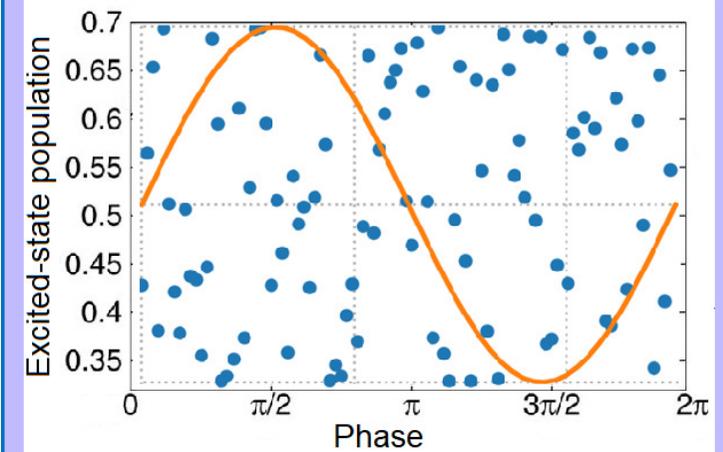


- Laser tilts, misalignments
- Laser phase noise, intensity noise
- Wavefront aberration
- Background magnetic fields
- ...



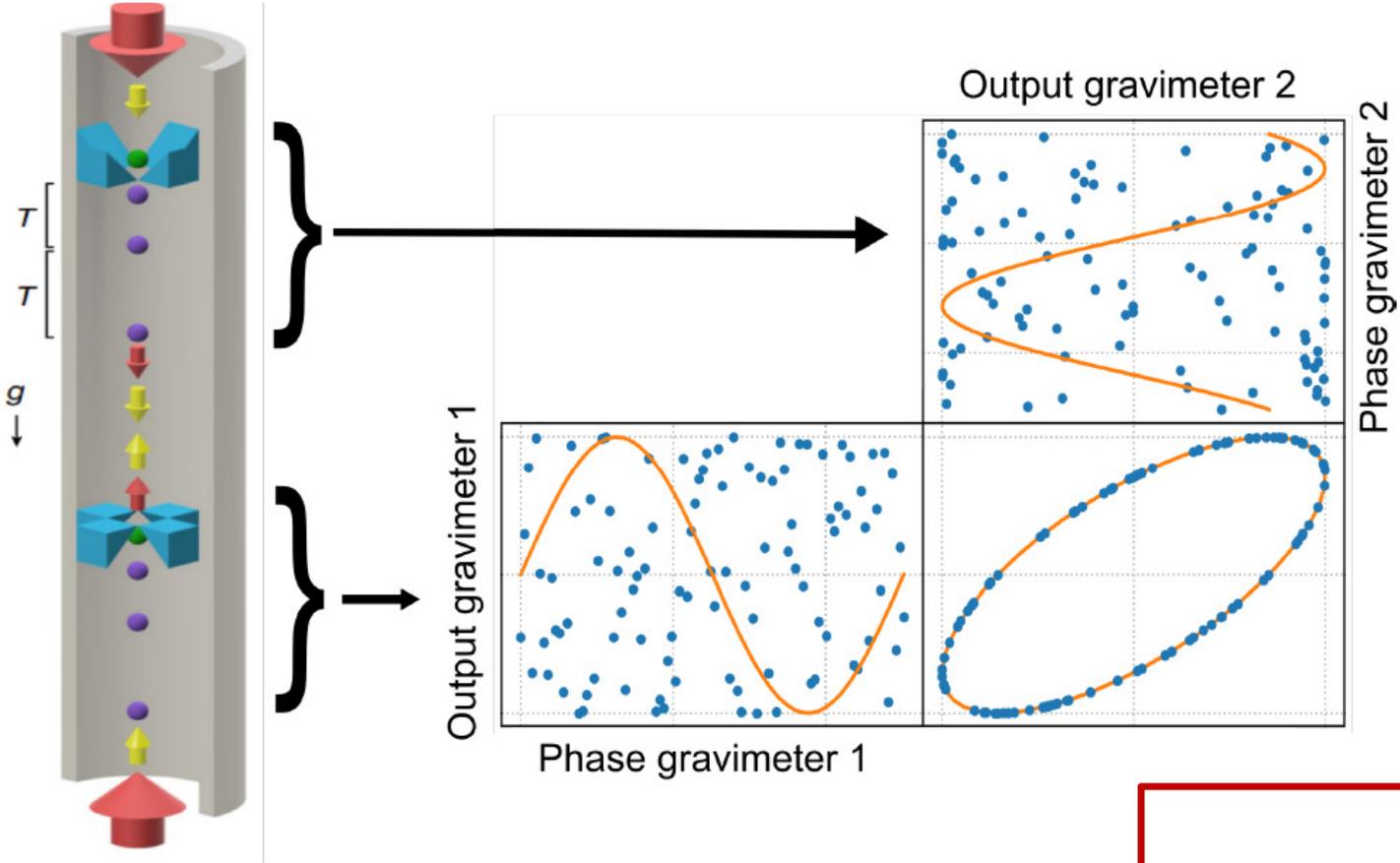
## External noise sources

### Unwanted additional contributions to the interferometric phase



- Vibrations
- Platform motion
- External field influences
- ...

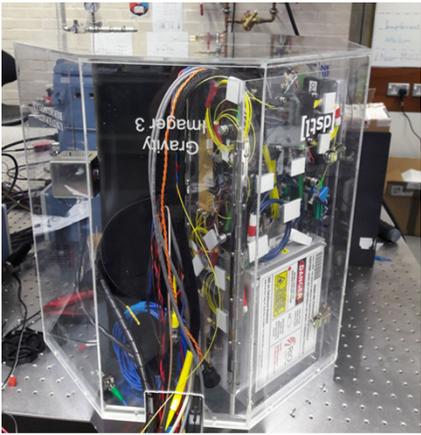
# Gravity gradiometry



$$\sigma_{\nabla g} = \frac{\sqrt{2}}{2Ck_{\text{eff}}T^2L\sqrt{N}}$$

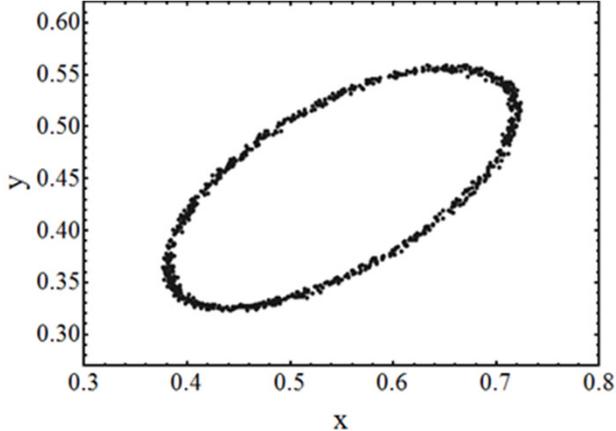
▲ Increase baseline

# Atom interferometry at UoB



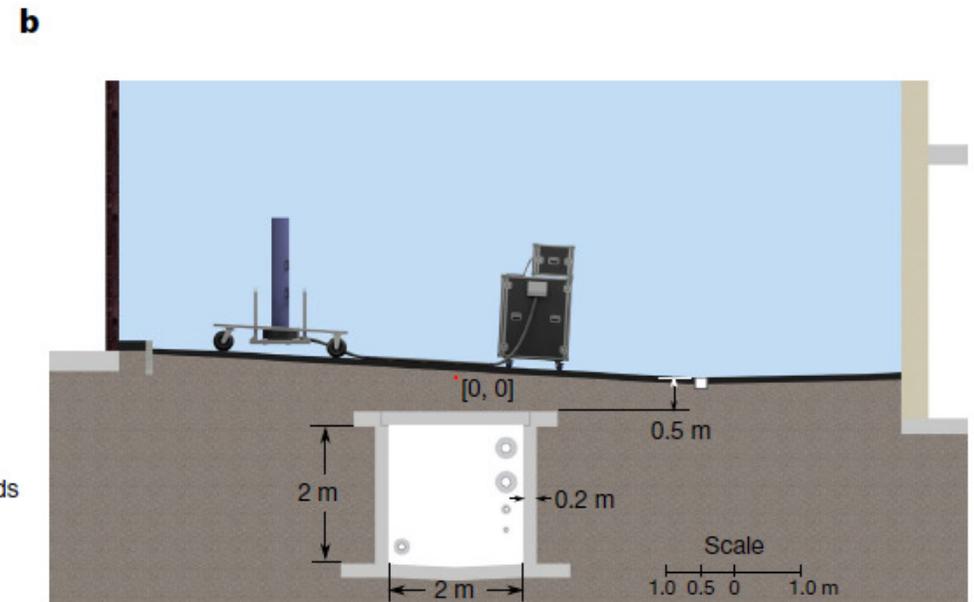
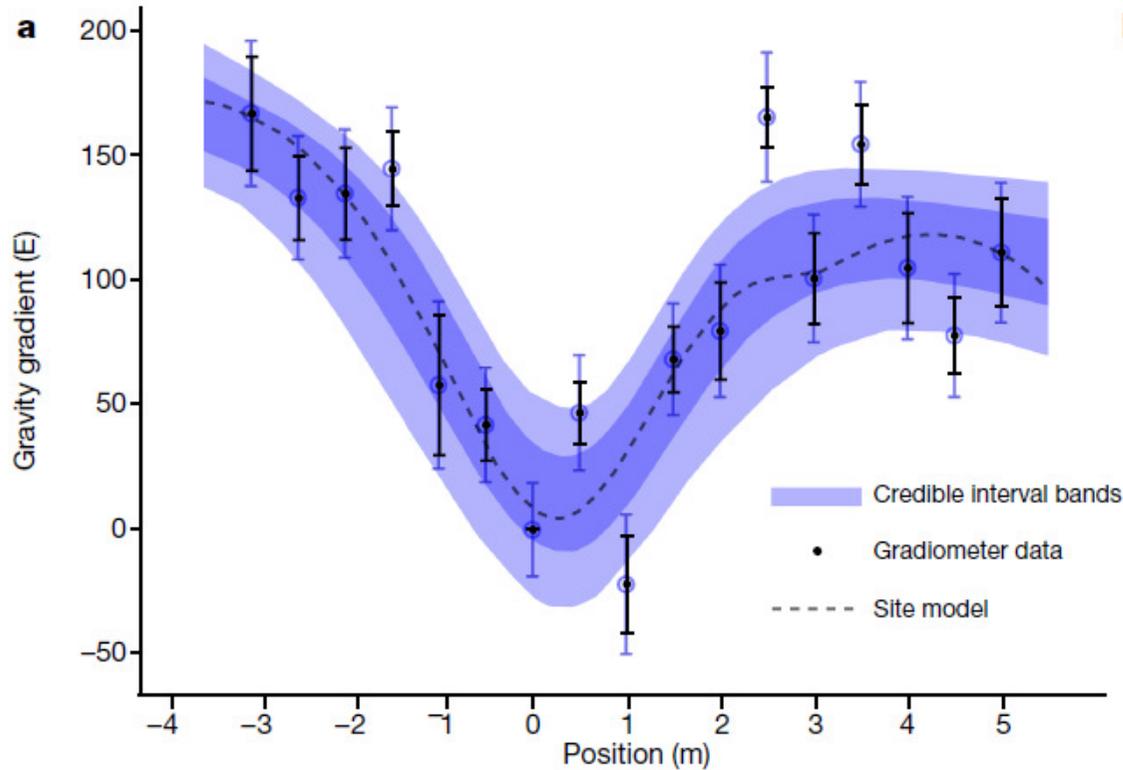
Top left: GGTOP. Top right: iSense project (EC collab.). Bottom: GI3 (DSTL).

Bottom: Field gravity gradiometer (DSTL, IUK, EPSRC). Right: Ellipse (11/2020)



# Atom interferometry in the field

Detection of a real-world application feature at UoB



B. Stray et al., *Quantum sensing for gravitational cartography*, *Nature* 602, 590–594 (2022)

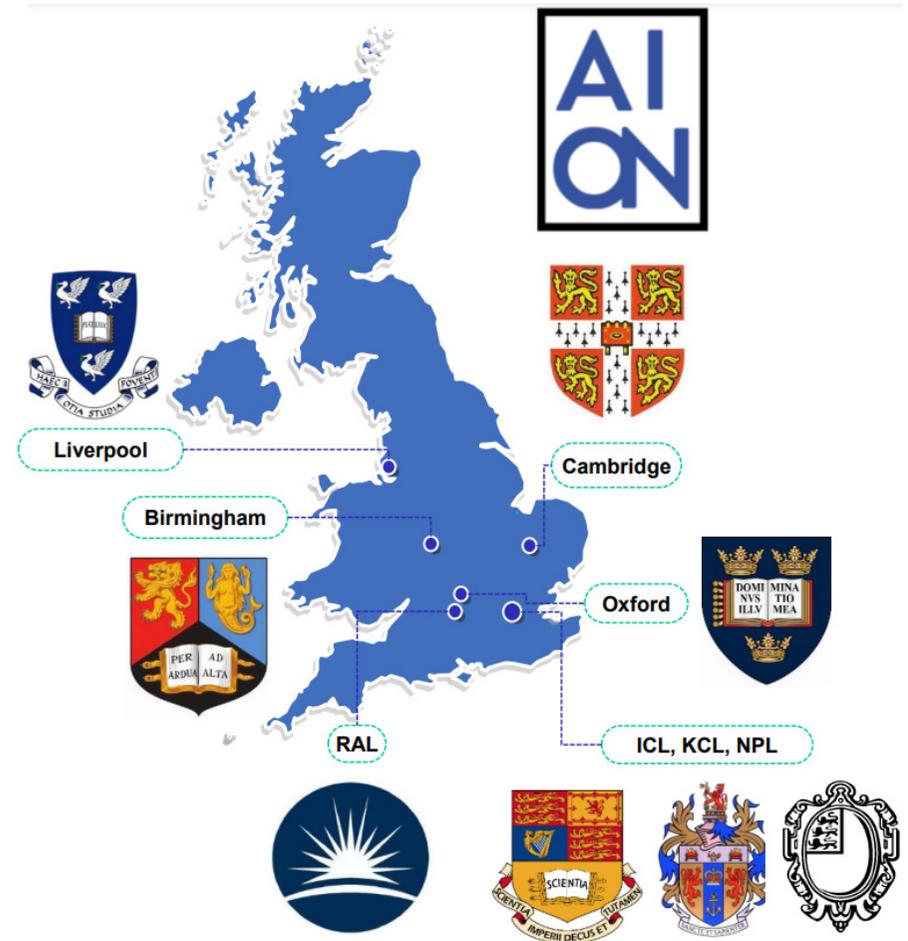
## **PART III.**

# **ATOM INTERFEROMETRY FOR FUNDAMENTAL PHYSICS: THE AION PROJECT**

# An Atom Interferometric Observatory and Network (AION)

Construct and operate an Atom Interferometric Observatory and Network (AION) that will :

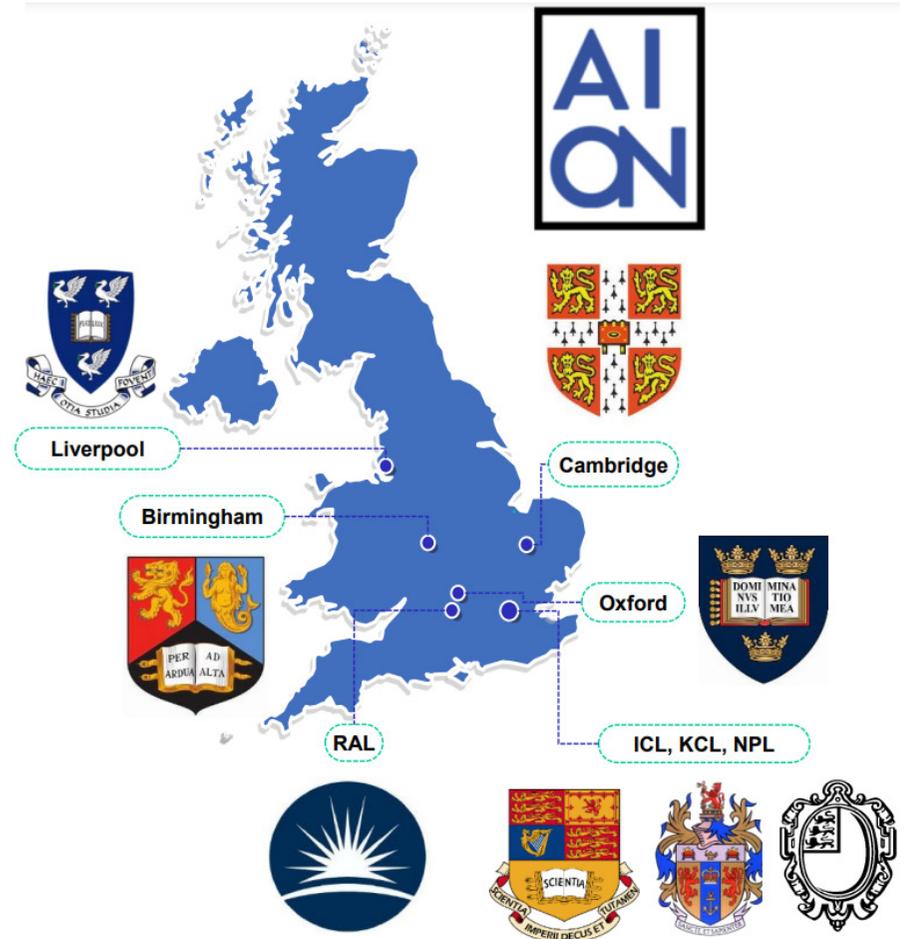
- enable the **exploration of properties of dark matter** and searches for new fundamental interactions.
- provide a pathway towards detecting **gravitational waves in the mid-frequency band** [0:01 Hz - few Hz] where currently operating and planned detectors are relatively insensitive.



# An Atom Interferometric Observatory and Network (AION)

Four stages corresponding to increased levels of performance:

- Stage 1: AION-10: 10m detector
- Stage 2: AION-100: 100m detector
- Stage 3: AION-1km: terrestrial 1km detector
- Stage 4: AION-SPACE: space detector



# Detecting gravitational waves (GW)

Resonant mass antennas (AURIGA, EXPLORER, ALLEGRO...)

Laser interferometers (LIGO, VIRGO, LISA...)

First detection of GW, 2015

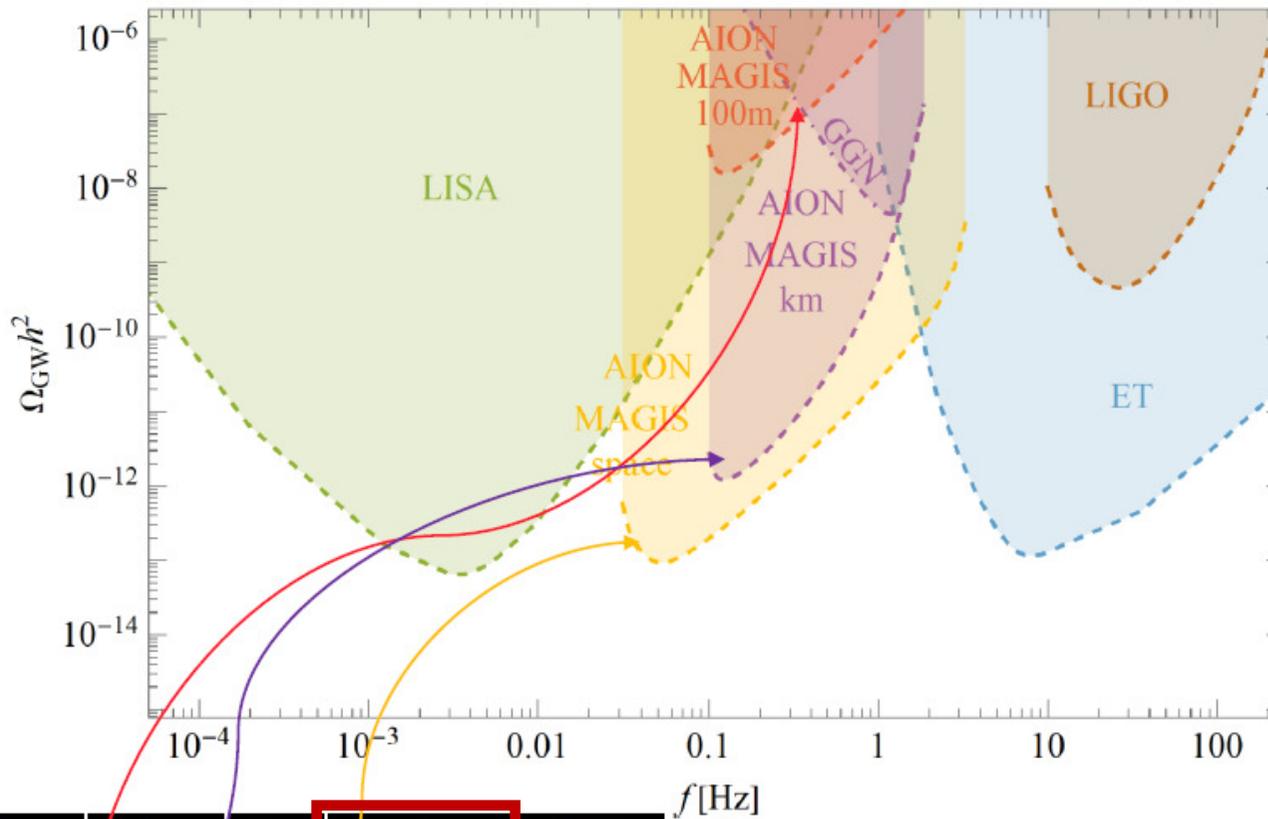
Remain limited by low-frequency noise  
in the mid-frequency band [0.01-10 Hz].



Atom interferometers (MAGIS, AION...)

A passing gravitational wave is detected via the strain it creates in the space between the free-falling atoms.

# Detecting mid-band gravitational waves



Sensitivity Scenario	L [m]	$T_{\text{int}}$ [s]	$\Phi$ [1/Hz]	LMP [#]
AION-100-today	100	1.4	$10^{-3}$	100
AION-100-ultimate	100	1.4	$10^{-5}$	40000
AION-km	2000	5	$0.3 \times 10^{-5}$	40000
AION-space	$4.4 \times 10^7$	300	$10^{-5}$	<1000

# Probing ultralight dark-matter (DM)

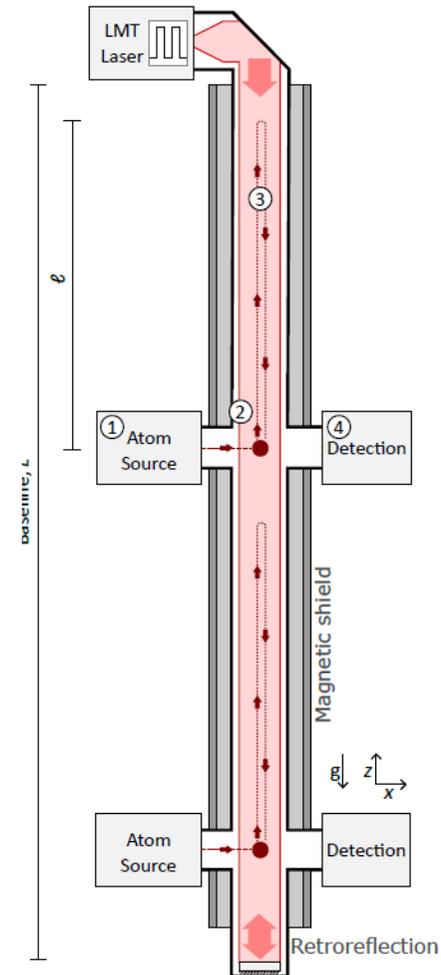
As of today, theoretical extensions of the standard model provide many candidate particles for DM, yet there has been no positive experimental result.



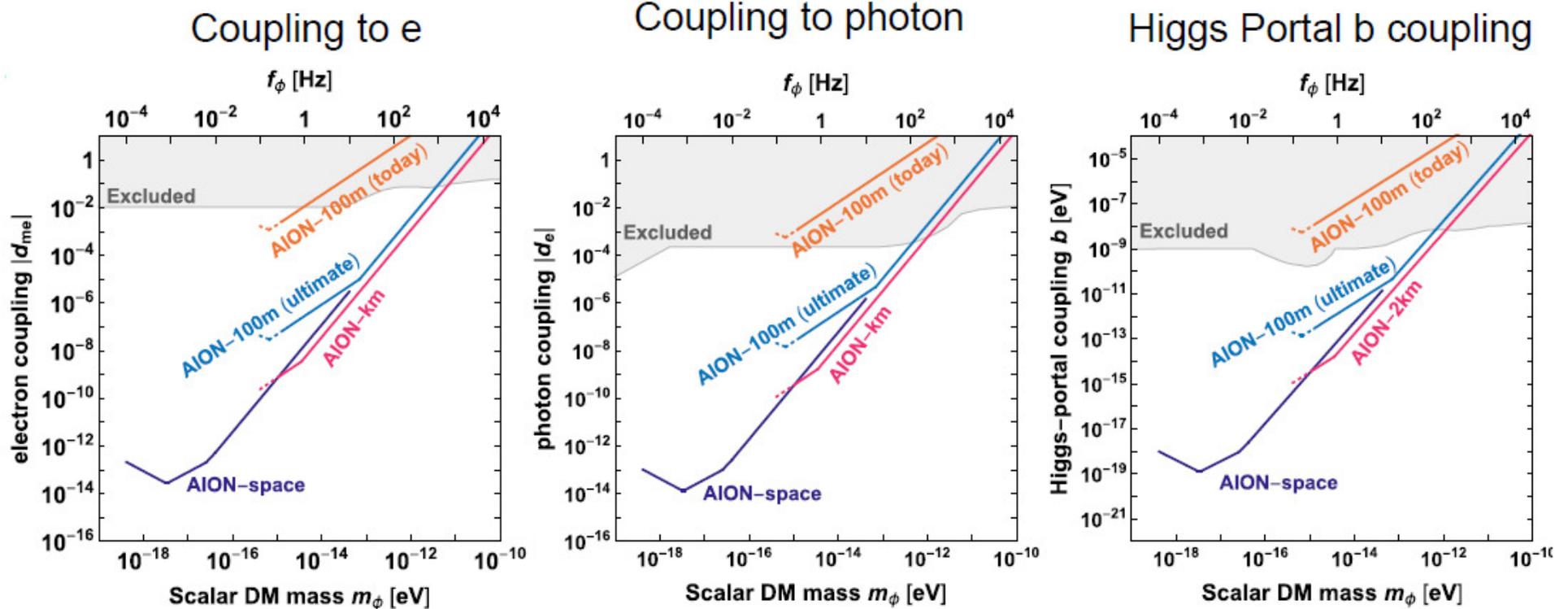
Reveal the nature of DM / blueprint a new method to probe the associated theoretical frameworks

Ultralight dark-matter induces a small time-dependent **perturbation to the atomic transition frequencies**.

Since the laser interacts with the separate interferometers at different times due to the light propagation delay, this perturbation will be observable as fluctuations in the differential phases accumulated by the separate interferometers.



# Probing ultralight dark-matter



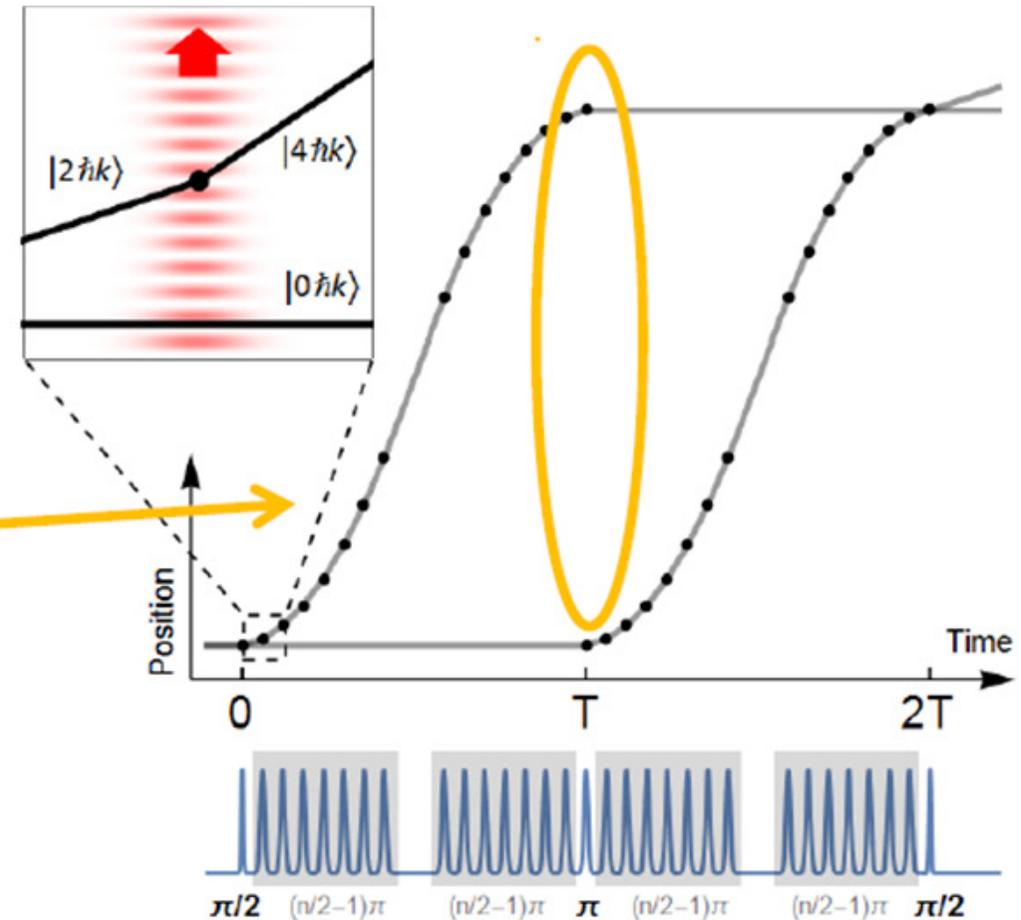
Based on: Arvanitaki et al., PRD **97**, 075020 (2018).

Sensitivity Scenario	L [m]	$T_{\text{int}}$ [s]	$\Phi$ [ $1/\sqrt{\text{Hz}}$ ]	LMP [#]
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# Large-momentum transfer (LMT) interferometry

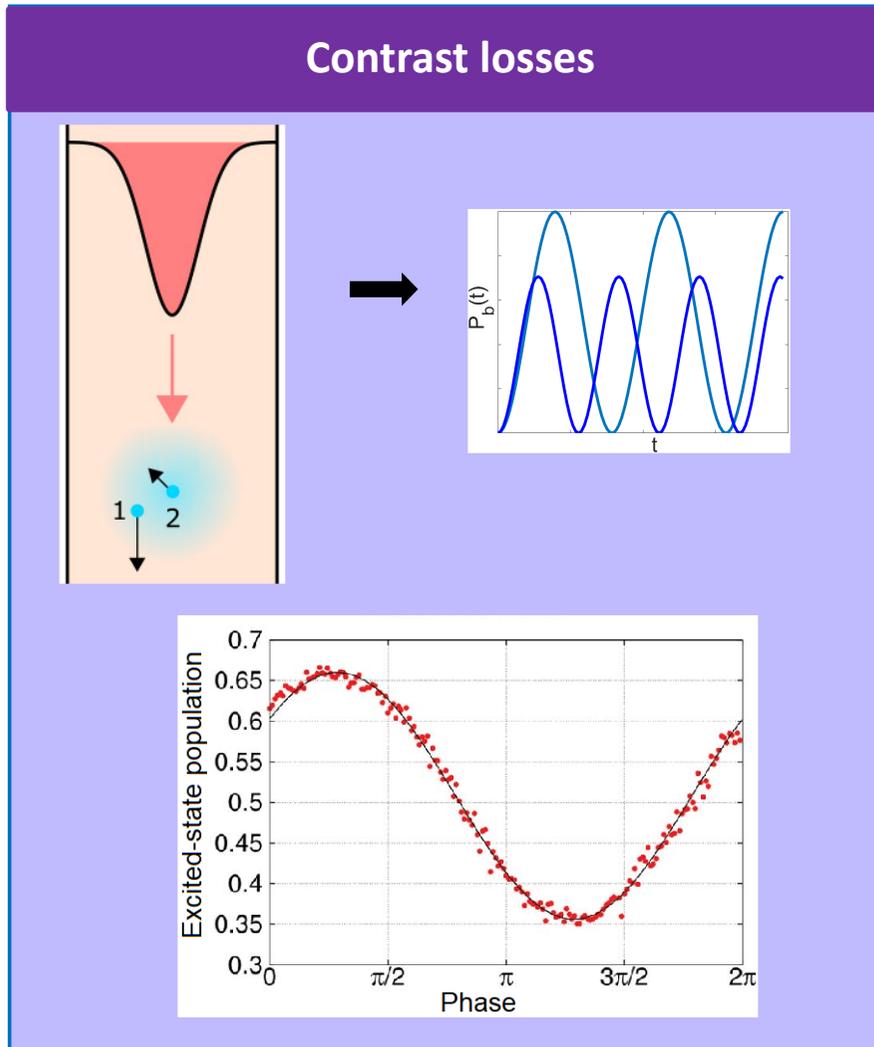
Increasing the momentum separation by applied successive light pulses

$N\hbar k$  momentum splitting



**Challenge: each light pulse must have a >99.9% efficiency !**

# Large-momentum transfer (LMT) interferometry

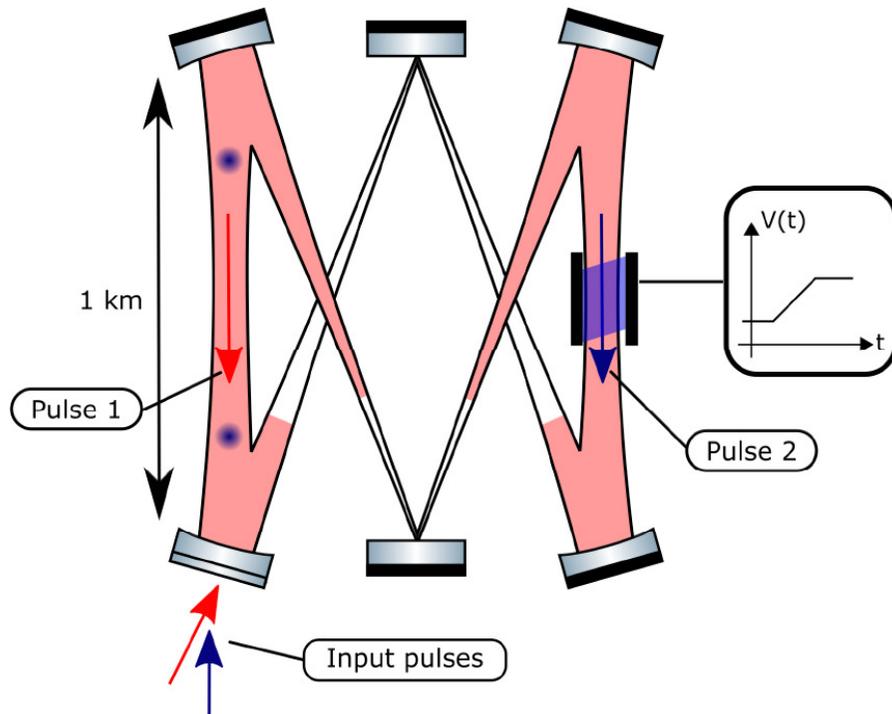


→ Very large beams/small clouds  
 Ultracold temperatures  
 High laser power

	AION 10 (goal)		AION 100 (goal)			
	689nm	698nm	689nm	698nm		
<b>Scheme</b>	689nm	698nm	689nm	698nm		
<b>Beam diameter (mm)</b>	10-20	10	-	50		
<b>Temperature (nK)</b>	0.1 - 1	1	0.1	-	1	0.1
<b>Laser power (W)</b>	> 20	A few 100	A few 1	-	135k	5k

# Enhanced atom-optics

## Cavity-based interferometry

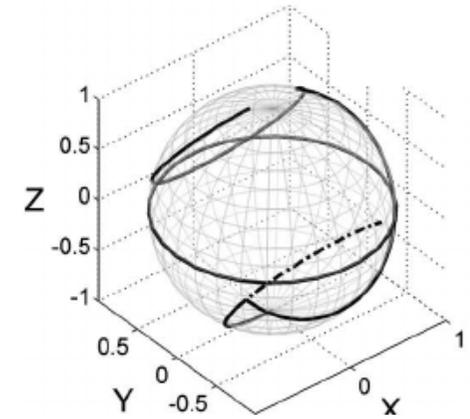


R. Nourshargh, S. Lellouch, S. Hedges, M. Langlois, K. Bongs, M. Holynski, Circulating pulse cavity enhancement as a method for extreme momentum transfer atom interferometry. *Communications Physics* 4, 257 (2021).

## Advanced pulse techniques

### Composite pulses

**Experimental trials at UoB**



### Floquet atom-optics

Wilkason T, Nantel M, Rudolph J, et al. Atom interferometry with floquet atom optics. *Phys Rev Lett.* 2022;129:183202.

### Shaped pulses

Lellouch et al. *EPJ Quantum Technology*  
<https://doi.org/10.1140/epjqt/s40507-023-00165-2>

**THANK YOU!**